

Supernovae as Sources of Interstellar Dust

Takaya Nozawa

Kavli IPMU, University of Tokyo

Main Collaborators:

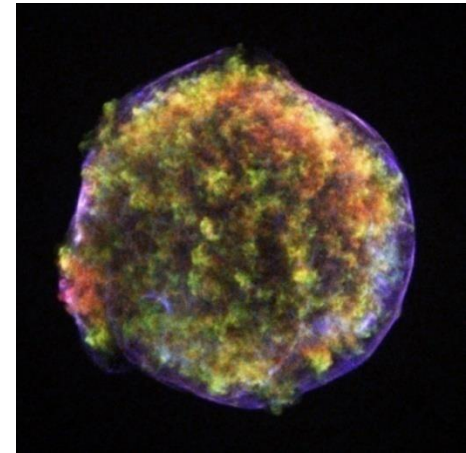
T. Kozasa (Hokkaido Univ.)

K. Maeda (IPMU)

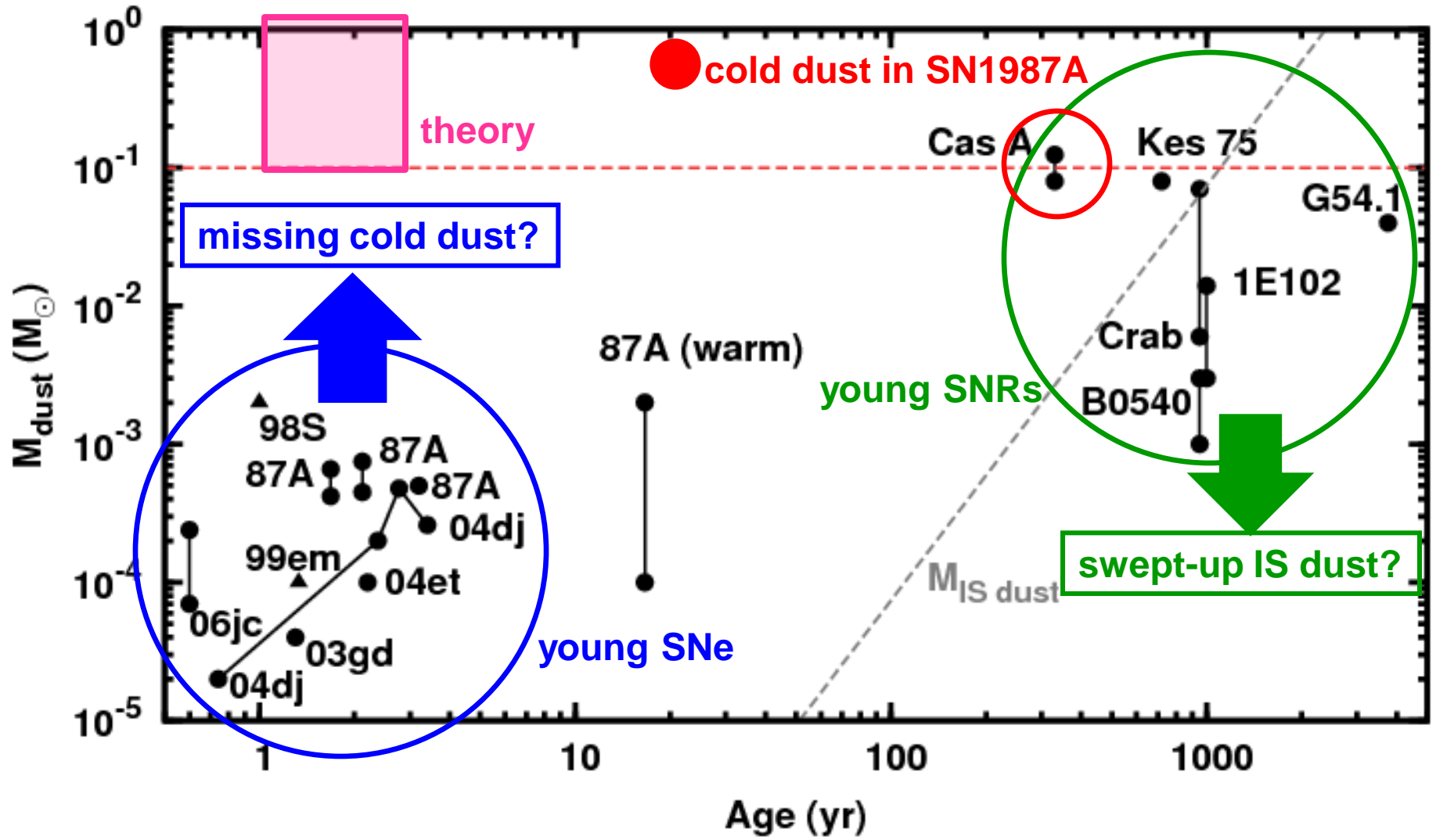
M. Tanaka (NAOJ)

K. Nomoto (IPMU)

H. Umeda (Univ. of Tokyo)



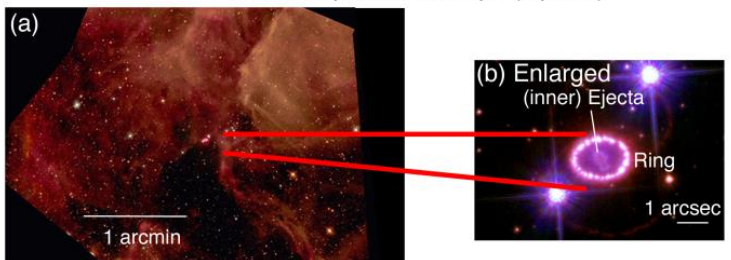
1. Missing-dust problem in core-collapse SNe



There is a big difference in the mass of dust between observational estimates and theoretical predictions!!

2. Herschel detects cold dust in SN 1987A

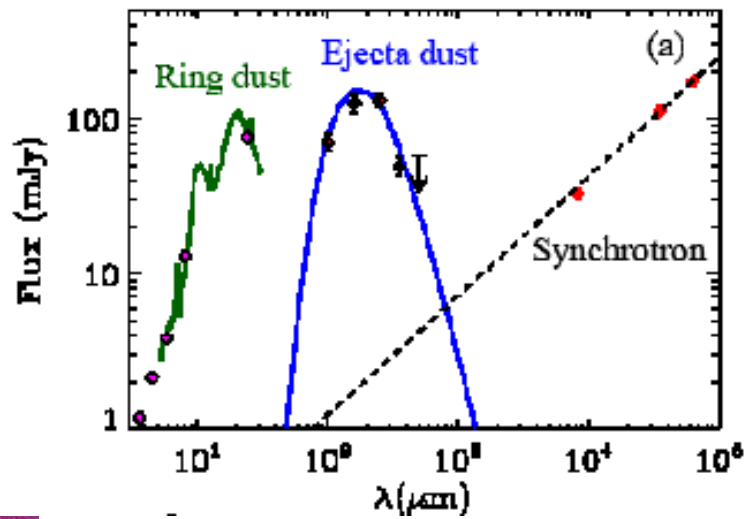
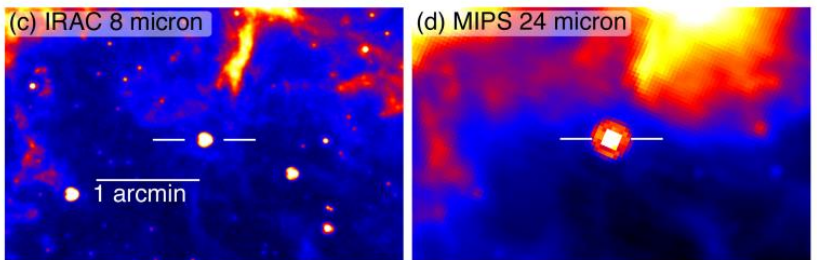
Hubble Space Telescope (Optical)



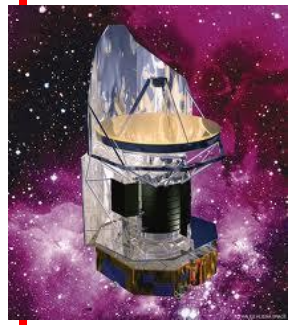
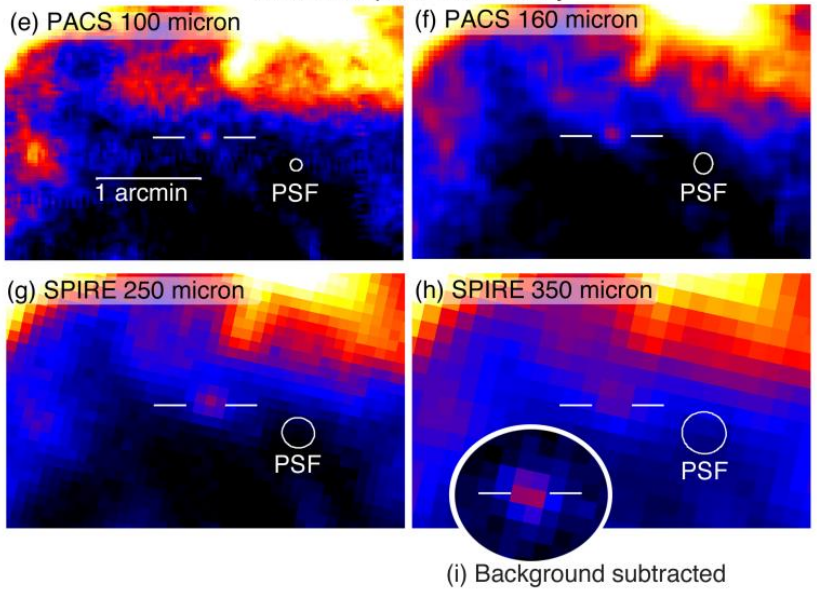
Matsuura, ... TN, et al. 2011



Spitzer Space Telescope



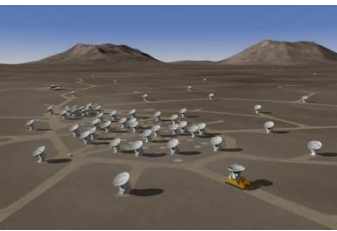
Herschel Space Observatory



Herschel detects cold (~20K) dust of ~0.4-0.7 Msun toward SN 1987A!

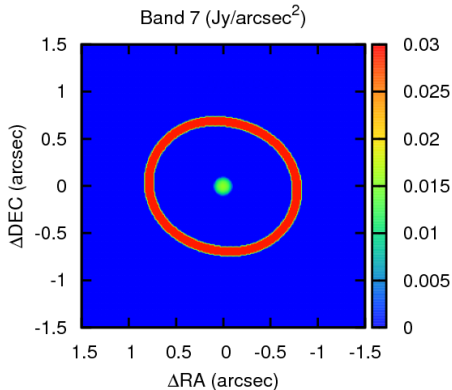
3. Resolving cold dust in SN 1987A with ALMA

ALMA Cycle 0 Proposal
'Detecting cool dust in SN1987A'
(TN, Tanaka, Moriya et al.)

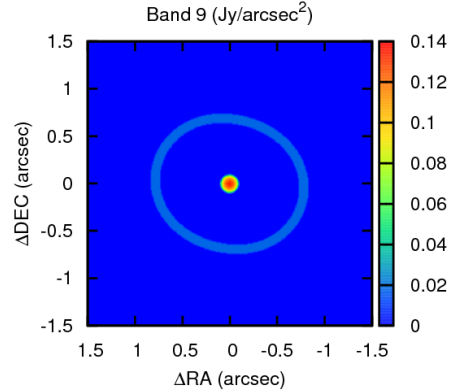


2 arcsec

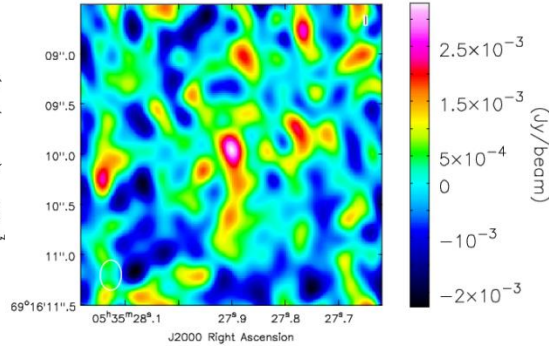
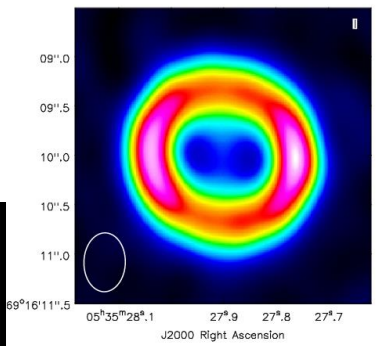
Band 7 (850 μm)



Band 9 (450 μm)



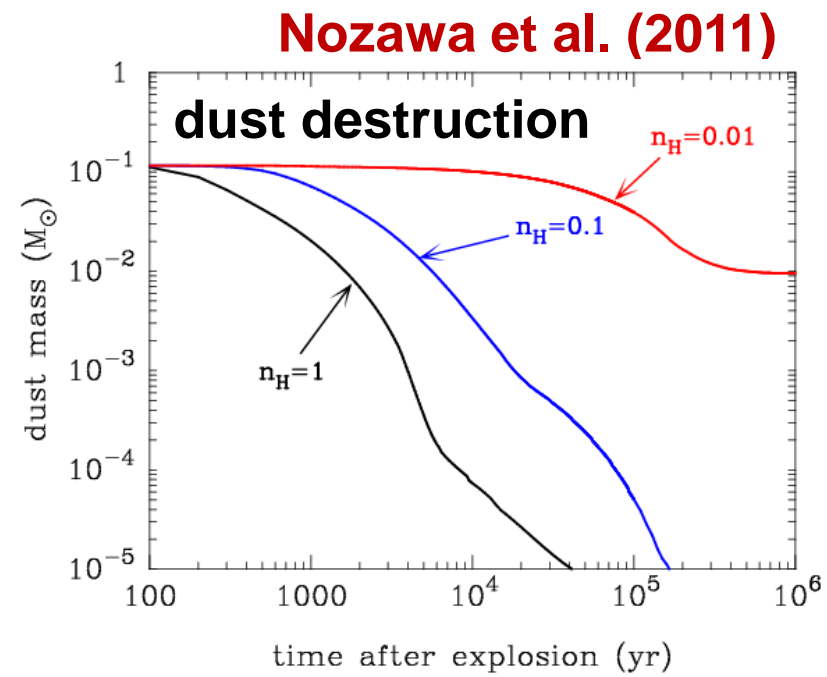
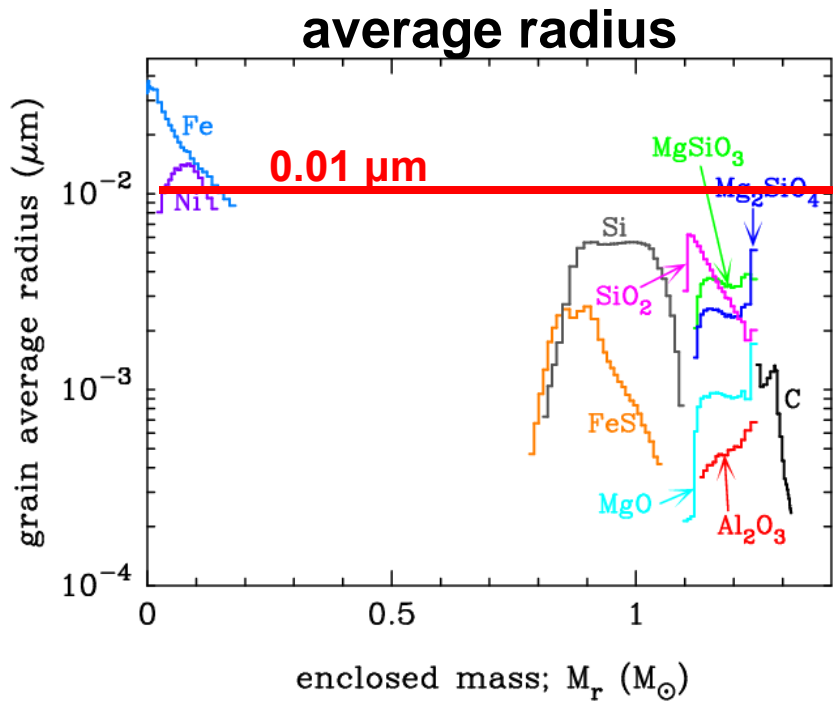
CASA simulation with extended configuration



This proposal was ranked in the highest priority, to be observed this summer

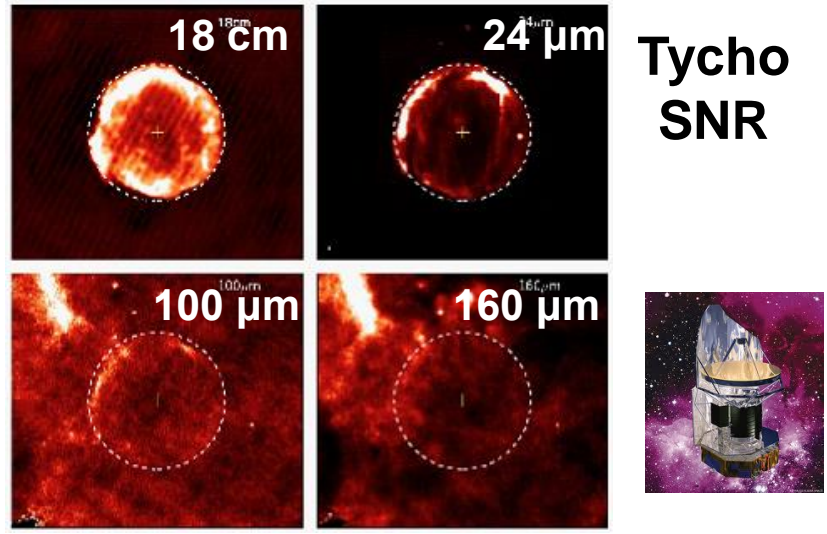
0.1 M_{sun} of silicate
→ 5 σ detection at Band 9 !!

4. Dust formation and evolution in SNe Ia



- average radius of dust : $a_{\text{ave}} \sim 0.01 \mu\text{m}$
- total dust mass : $M_{\text{dust}} = 0.1\text{-}0.2 M_{\text{sun}}$

SNe Ia are unlikely to be major sources of dust

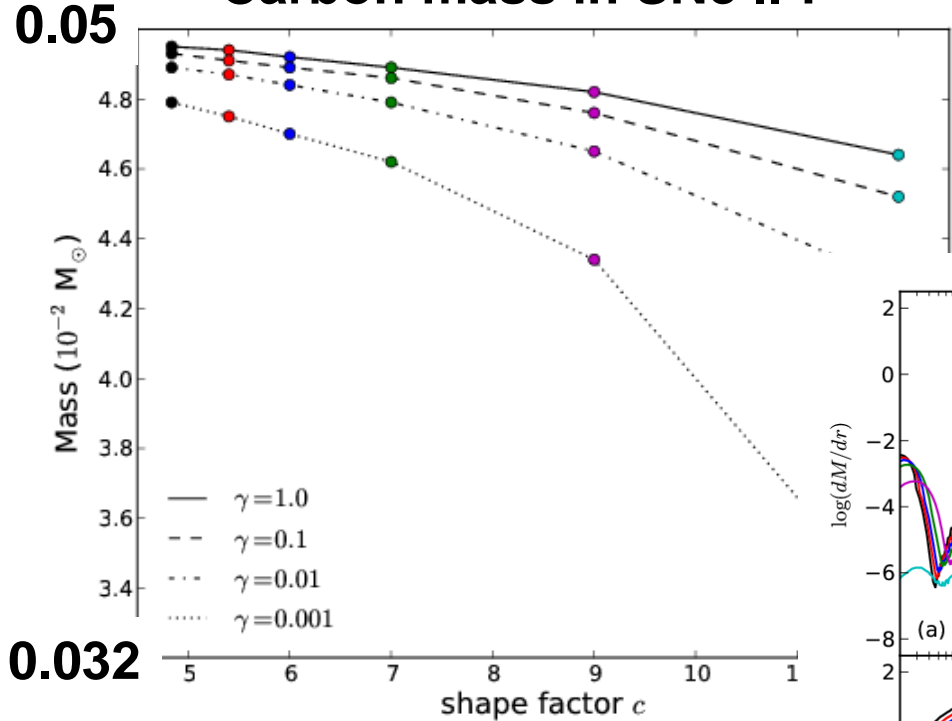


Gomez, Clark, TN, et al. (2011)

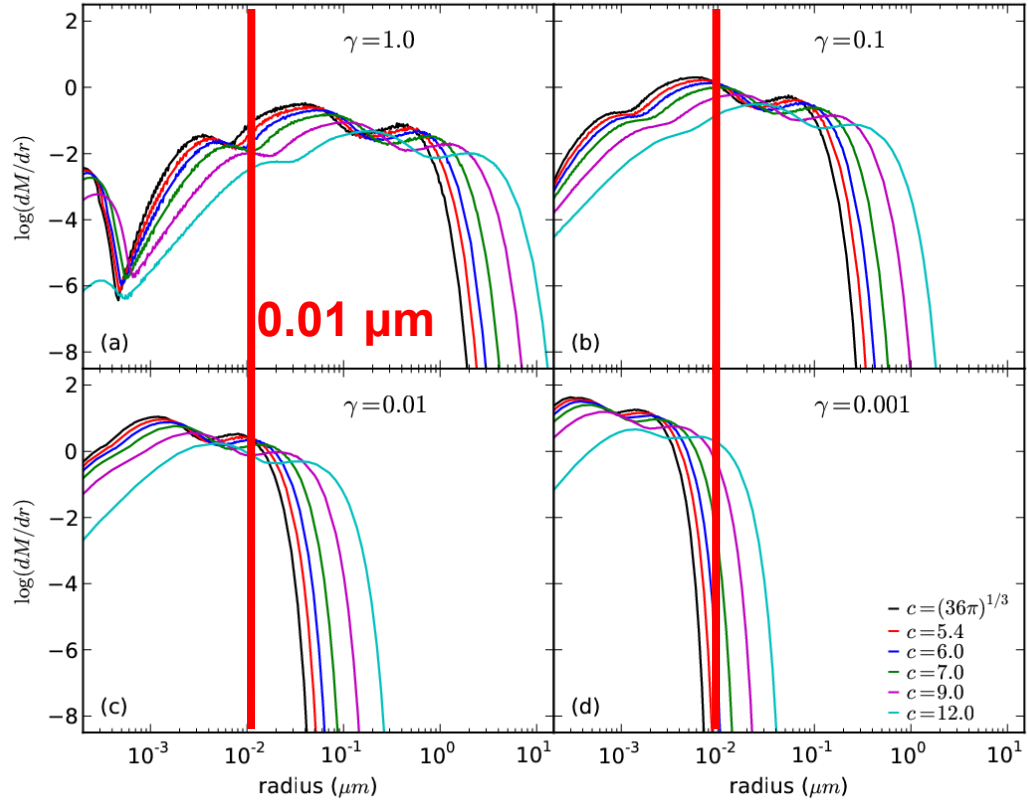
5. Effects of nano-properties on dust formation

Falless, TN, et al. (2011)

Carbon mass in SNe II-P



Size distribution of carbon grains



0.032 **spherical** \longrightarrow **asph**

Mass of carbon grains formed in SNe II-P is insensitive to sticking probability of ≥ 0.001