

# High Density Matter from QCD



<https://www.youtube.com/watch?v=MTvgnYGu9bg>

A YouTube video player interface showing a video titled "Muse - Neutron Star Collision (Love Is Forever) [OFFICIAL VIDEO]". The video has been playing for 2:06 of its 4:08 duration. The player includes standard controls like play/pause, volume, and settings.

Muse - Neutron Star Collision (Love Is Forever) [OFFICIAL VIDEO]

第27回 理論懇シンポジウム「理論天文学・宇宙物理学と境界領域」

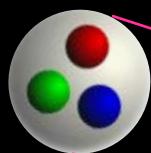
2014年 12月24日(水)

Tetsuo Hatsuda (RIKEN)

The progress of science requires the growth of understanding in both directions, downward from the whole to the parts and upward from the parts to the whole.

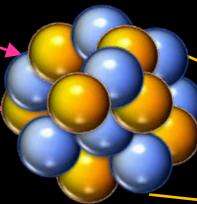
From "The Scientist as Rebel" by Freeman J. Dyson

nucleon



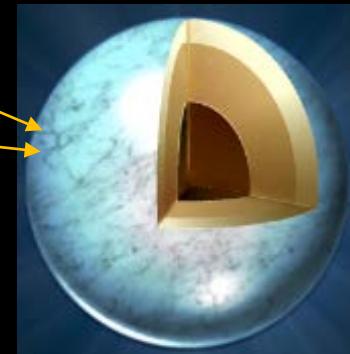
$r \sim 1 \text{ [fm]}$

nucleus

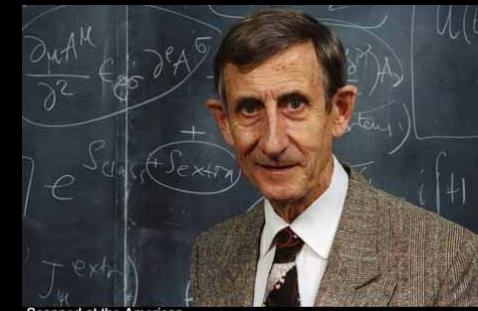


$r \sim 10 \text{ [fm]}$

Neutron star



$r \sim 10 \text{ [km]}$



Scanned at the American Institute of Physics

# Plan of this Talk

1. Introduction : phase structure of QCD
2. Baryon forces from Lattice QCD (LQCD)

Inoue et al. [HAL QCD Coll.], PRL 106 (2011) 162002

3. Nuclear & Neutron Matter from LQCD

Inoue et al. [HAL QCD Coll.], PRL 111 (2013) 112503

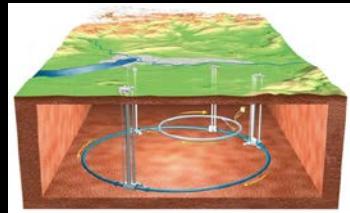
4. Finite Nuclei from LQCD

Inoue et al. [HAL QCD Coll.], arXiv: 1408.4892

5. Summary

# QCD Phase Structure

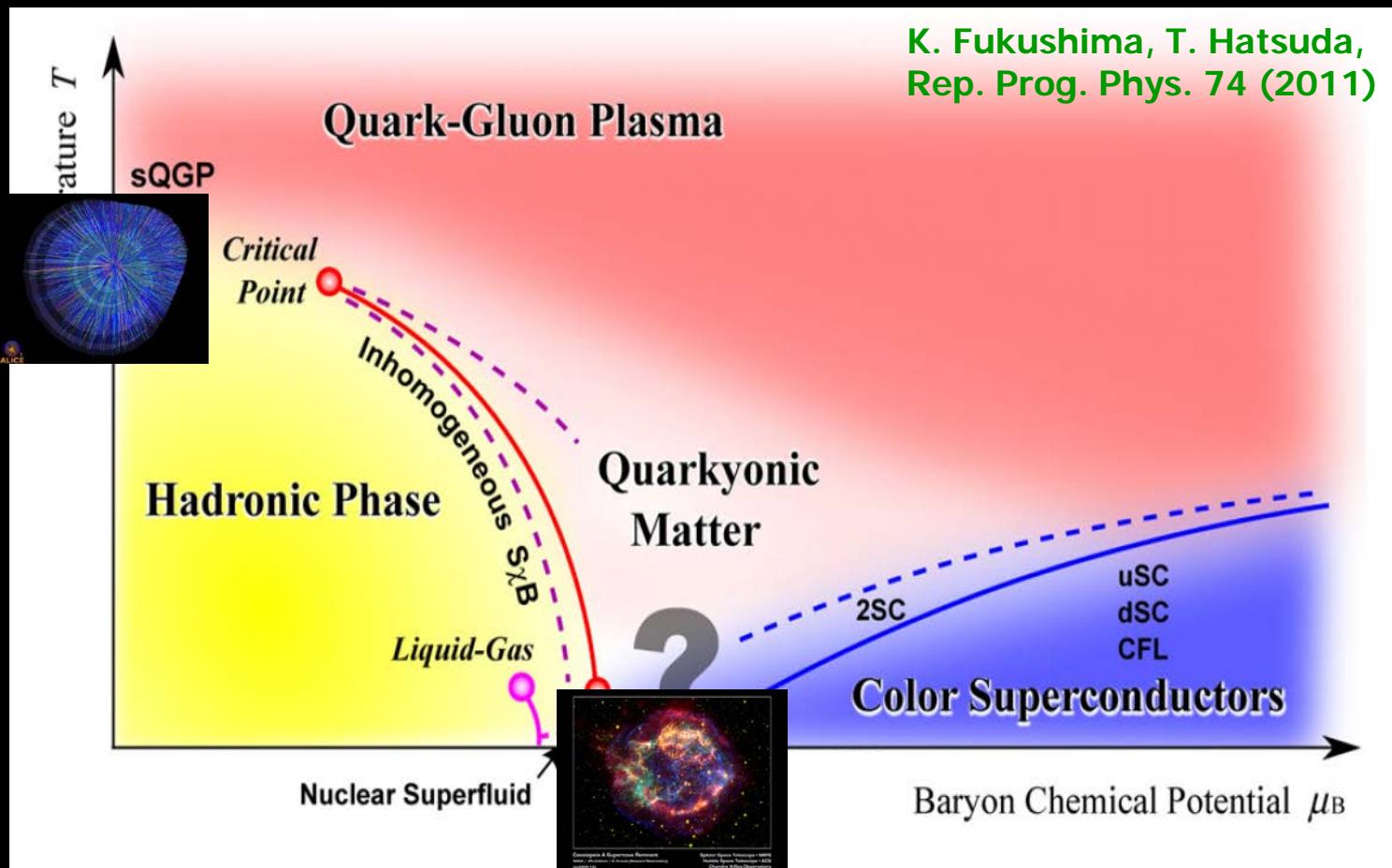
Physics of  
Early Universe



LHC @ CERN



RHIC @ BNL



J-PARC



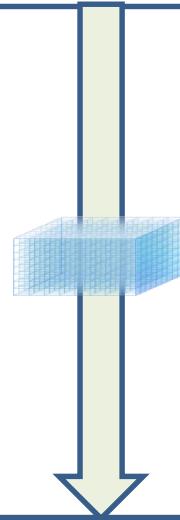
KAGRA

Physics of  
Neutron stars  
and Black holes

# From QCD to Hot Matter

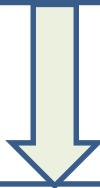
## Quantum Chromo Dynamics

Lattice QCD

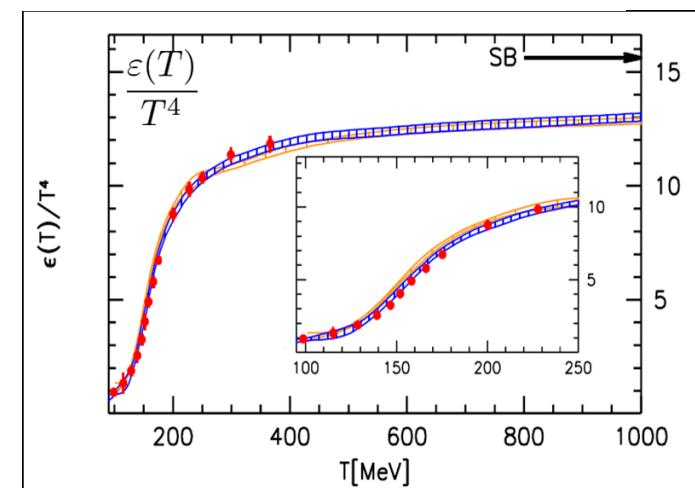
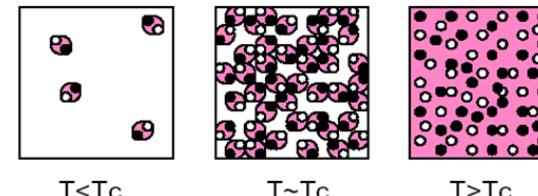


Equation of State for Hot Matter

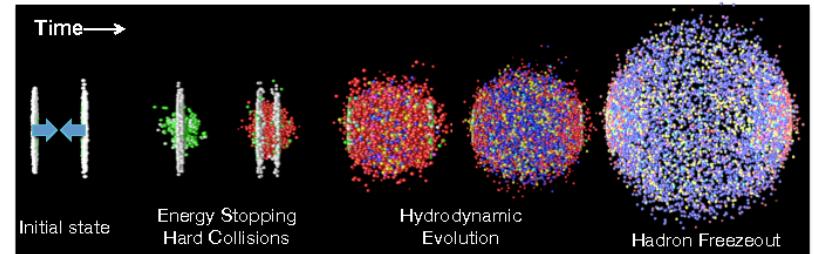
Relativistic  
hydrodynamics



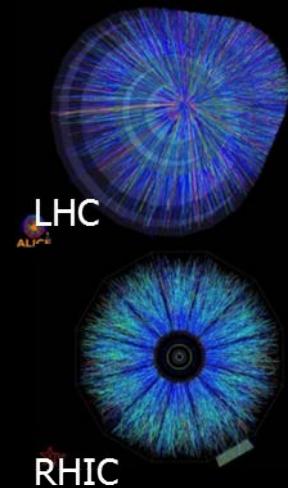
Relativistic heavy-ion collisions



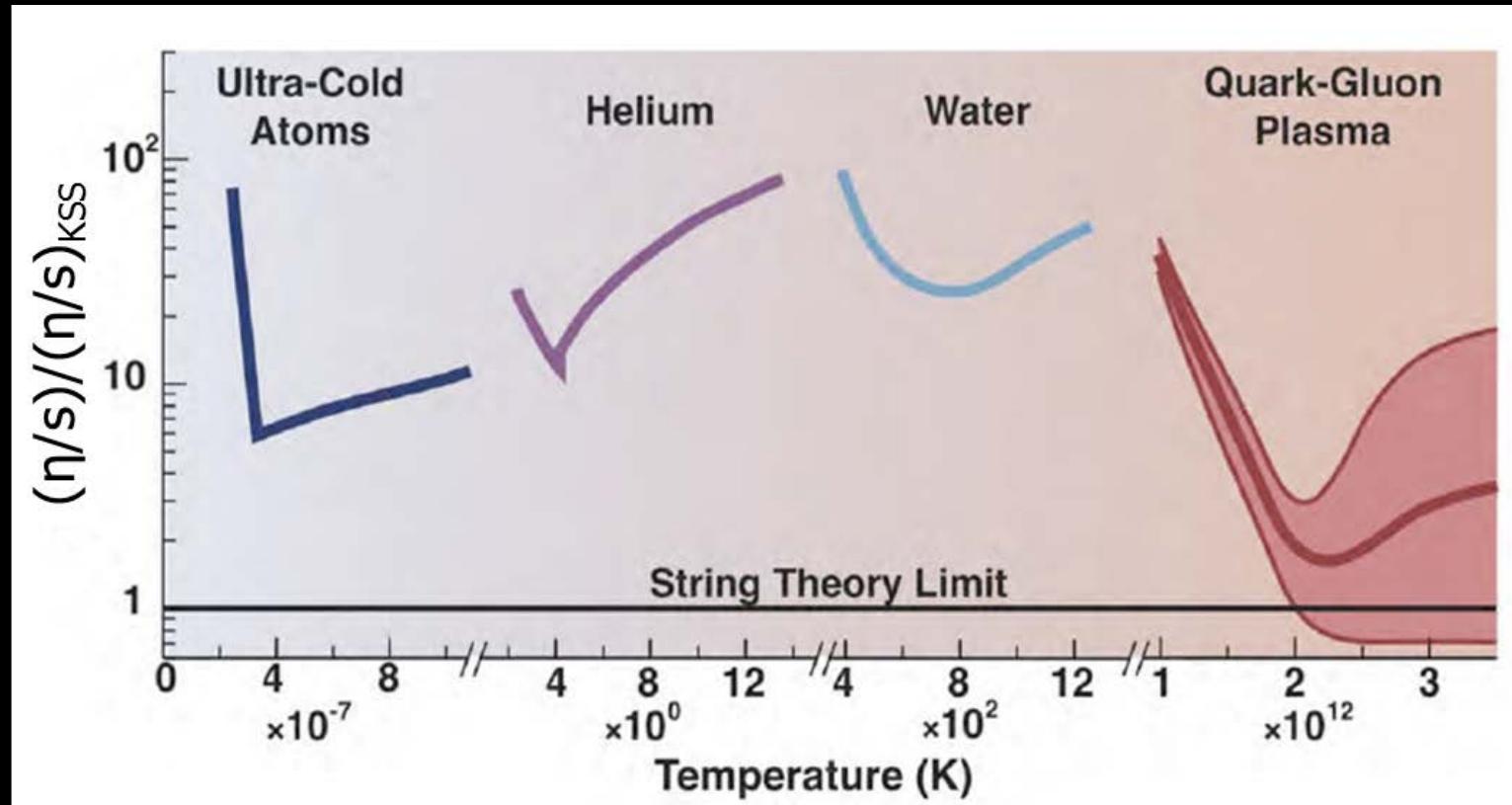
Wuppertal-Budapest Coll. JHEP 1011 (2010) 77



# Quark-Gluon Plasma at RHIC/LHC



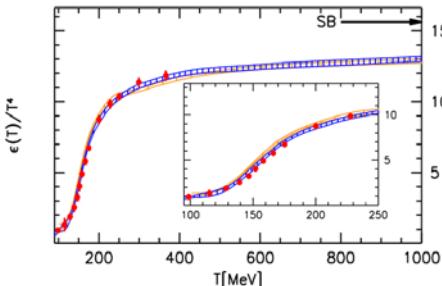
- high temp.:  $T \sim 4 \times 10^{12} \text{ K}$
- low viscosity:  $\eta/s \sim 1/4\pi$



# From QCD to Dense Matter

## Quantum Chromo Dynamics

Lattice gauge simulations



sign problem



Phenomenological nuclear force

Equation of State for Hot Matter

Relativistic hydrodynamics

Relativistic heavy-Ion collisions

Baryon interactions

Many-body techniques

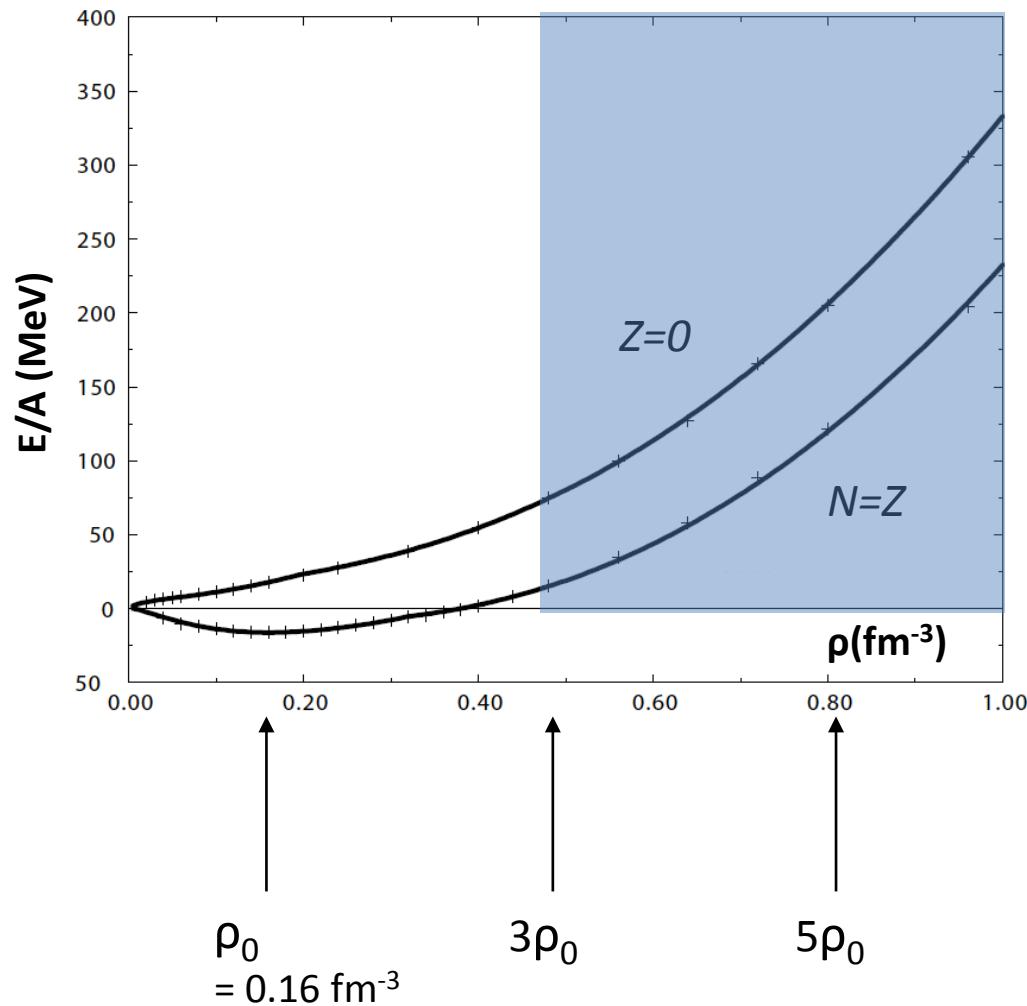
Equation of State for Dense Matter

General relativity

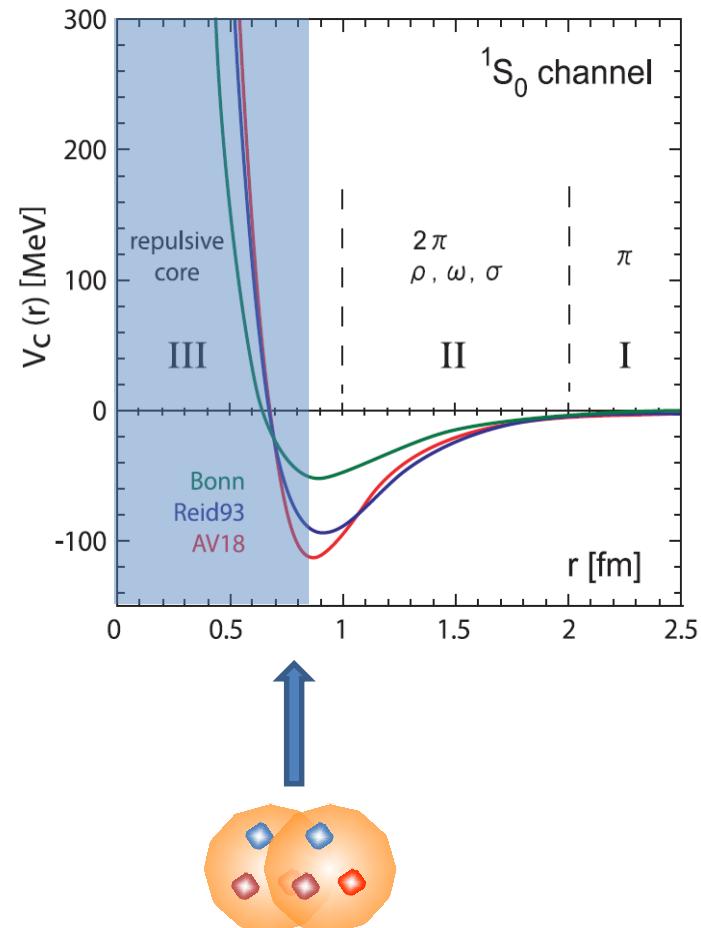
Neutron stars

# Nuclear Force and dense EOS (nucleons only)

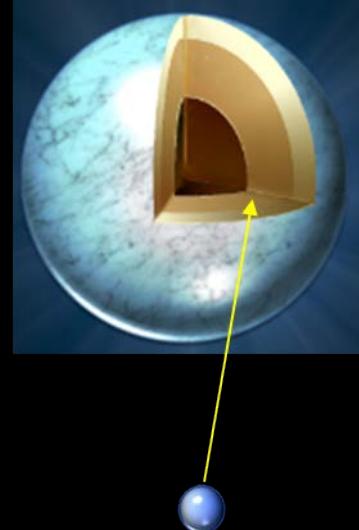
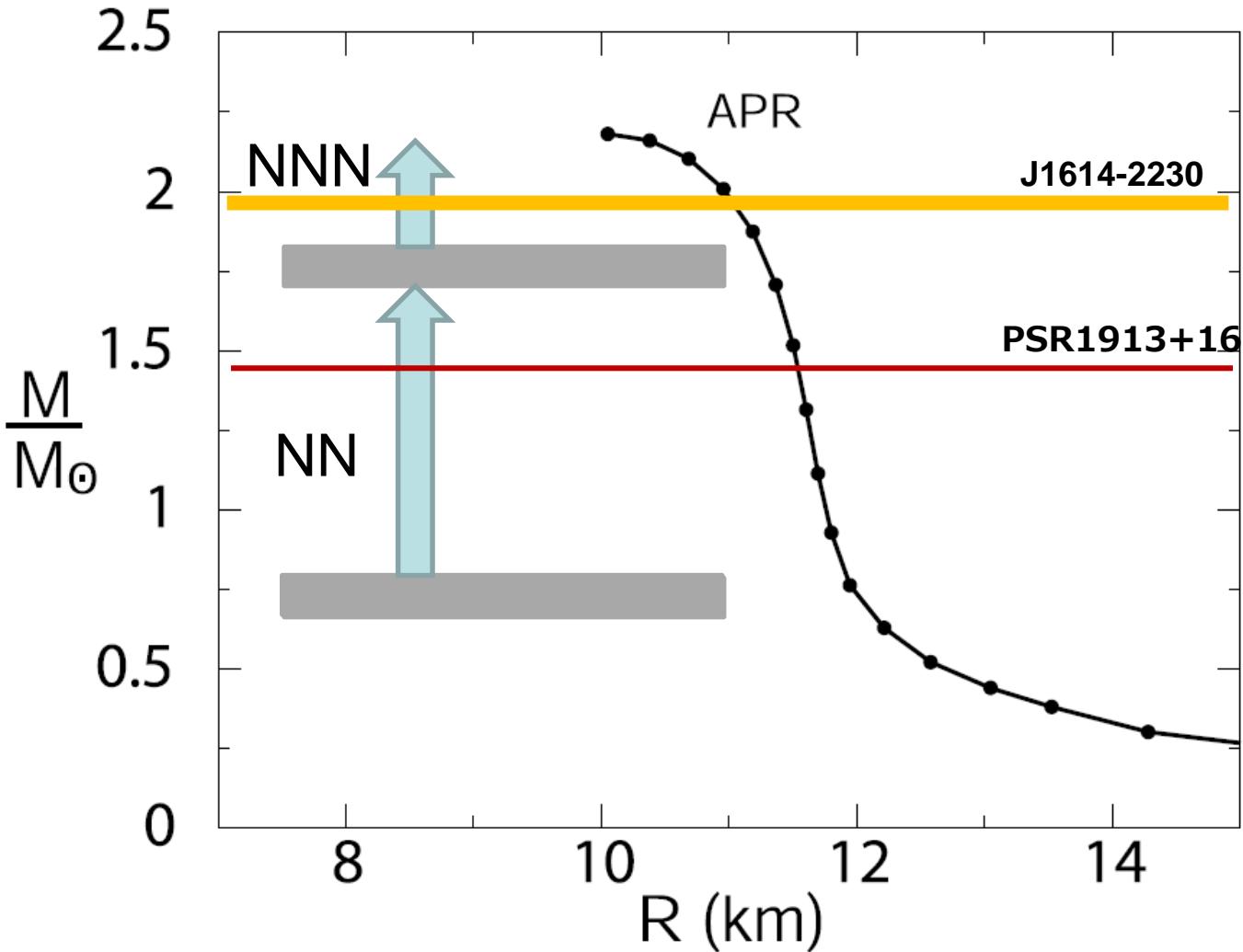
Akmal, Pandharipande & Ravenhall, PRC58 ('98)



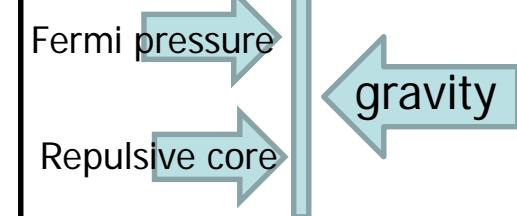
Phenomenological nuclear force



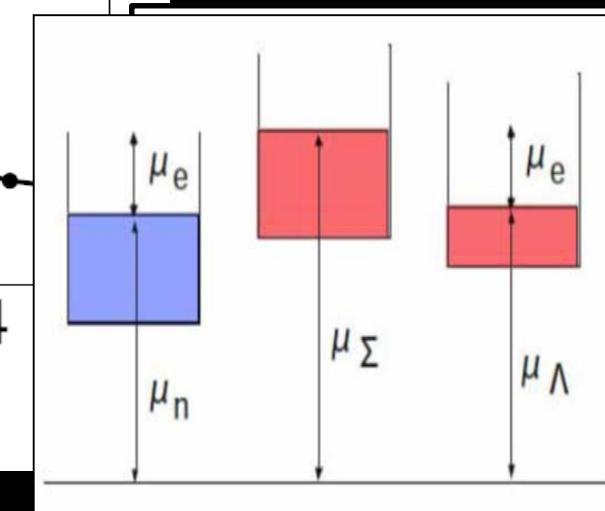
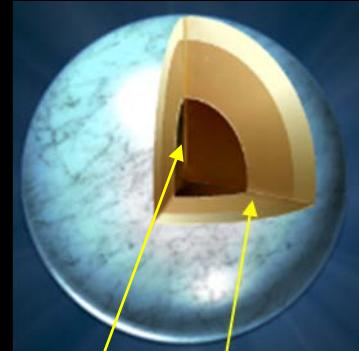
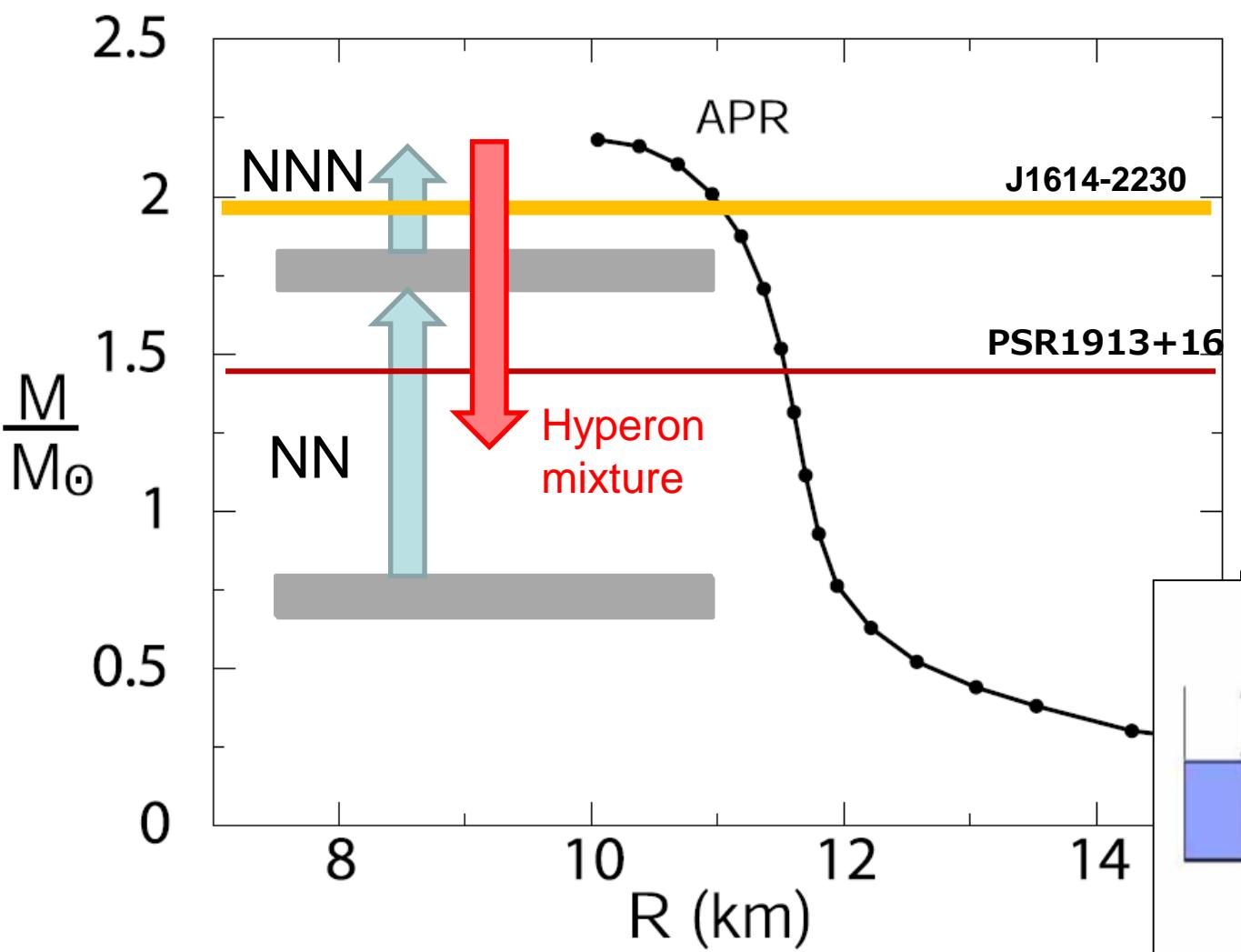
# Mass-Radius relation of $N_{\star}$ (nucleons only)



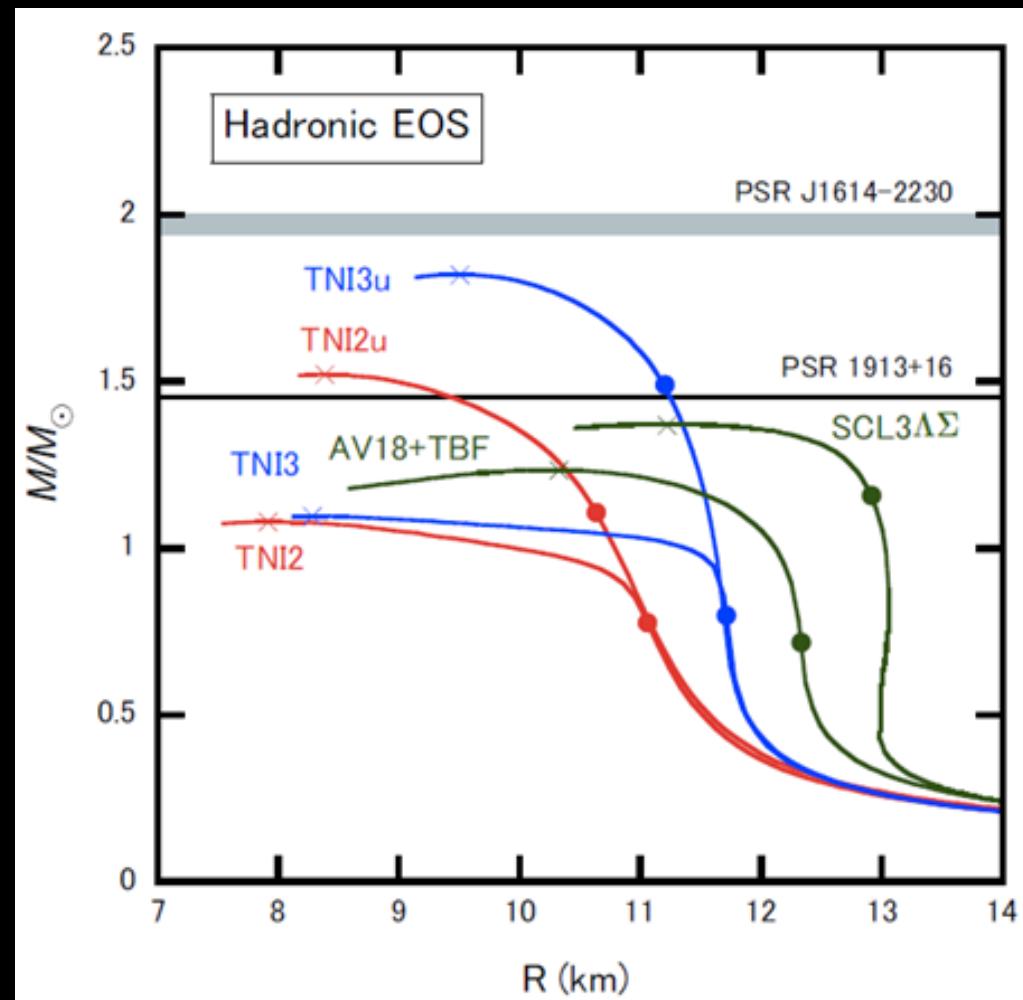
Pressure balance



# Hyperon Crisis (Takatsuka et al., 2002)



# Hyperon Crisis

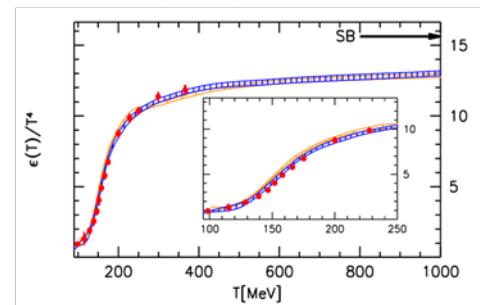


Masuda, Hatsuda & Takatsuka,  
Astrophysical Journal Letters 764 (2013) 12

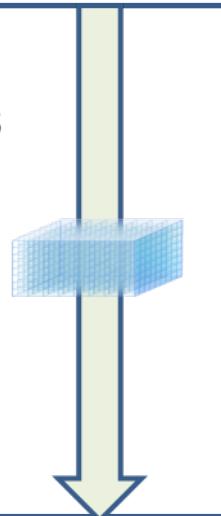
# From QCD to Hot/Dense Matter

## Quantum Chromo Dynamics

Lattice gauge simulations



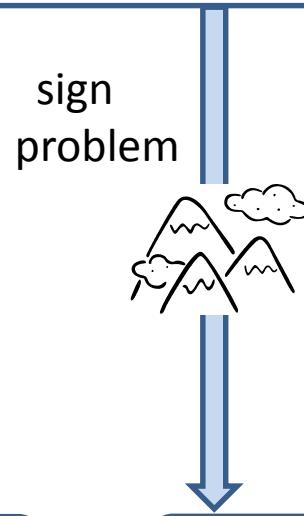
sign problem



Equation of State for Hot Matter

Relativistic hydrodynamics

Relativistic heavy-Ion collisions



Equation of State for Dense Matter

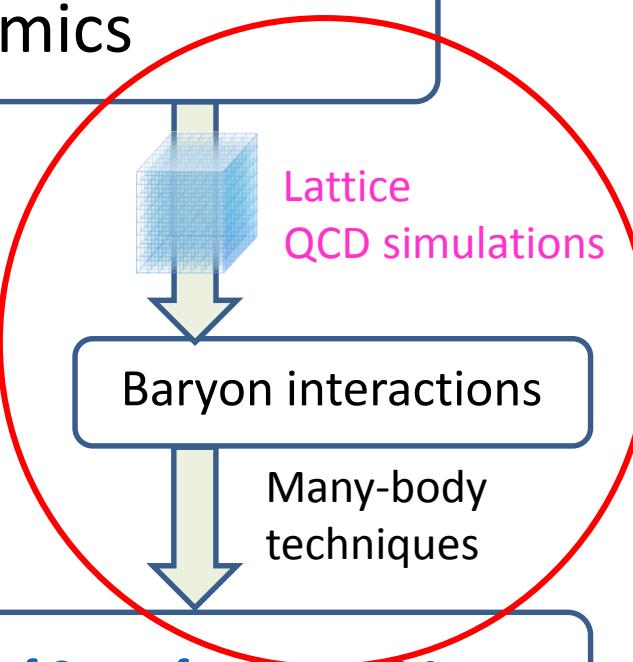
General relativity

Neutron stars

Lattice QCD simulations

Baryon interactions

Many-body techniques



# Precision QCD

$$\mathcal{L} = -\frac{1}{4}G_{\mu\nu}^a G_a^{\mu\nu} + \bar{q}\gamma^\mu(i\partial_\mu - g t^a A_\mu^a)q - m\bar{q}q$$

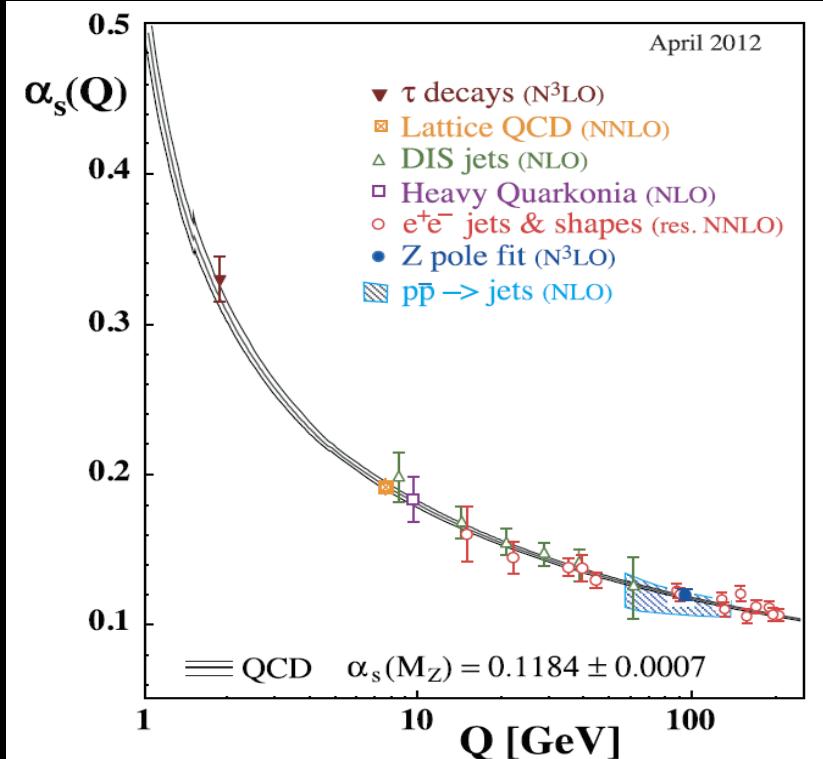
$$G_{\mu\nu}^a = \partial_\mu A_\nu^a - \partial_\nu A_\mu^a + g f_{abc} A_\mu^b A_\nu^c$$

Running masses:  $m_q(Q)$

| quark masses<br>(from lattice QCD) | [MeV]<br>(MS-bar @ 2GeV) |
|------------------------------------|--------------------------|
| $m_u$                              | 2.16 (9)(7)              |
| $m_d$                              | 4.68 (14)(7)             |
| $m_s$                              | 93.8 (2.4)               |

FLAG Collaboration update( July 26, 2013)  
<http://itpwiki.unibe.ch/flag/>

Running coupling:  $\alpha_s(Q)=g^2/4\pi$

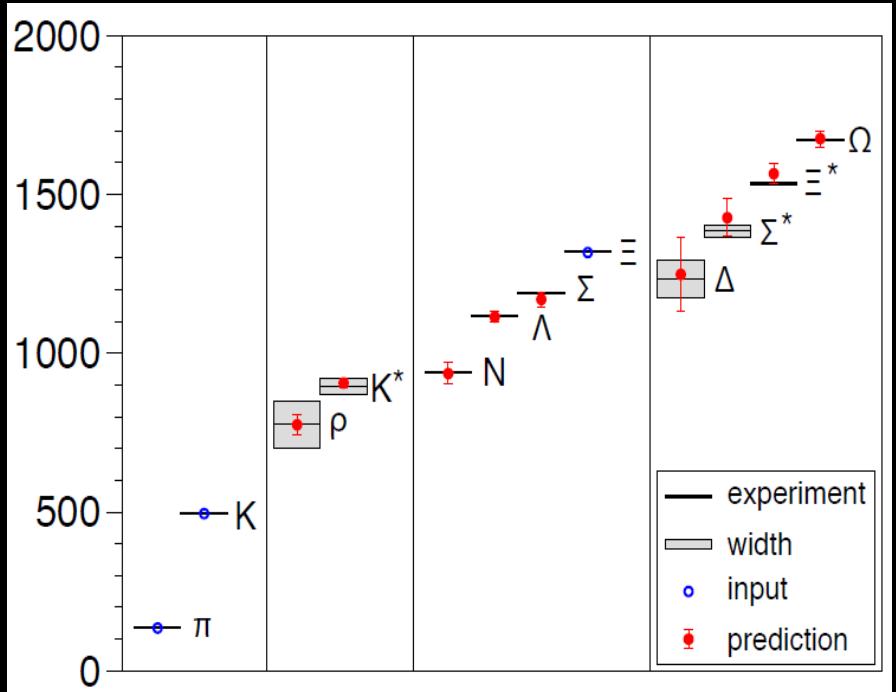
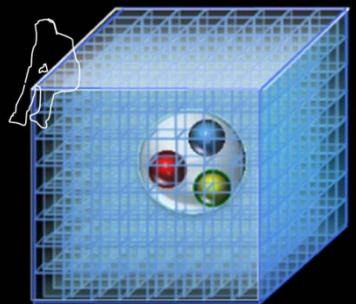


# Hadron masses (2008)

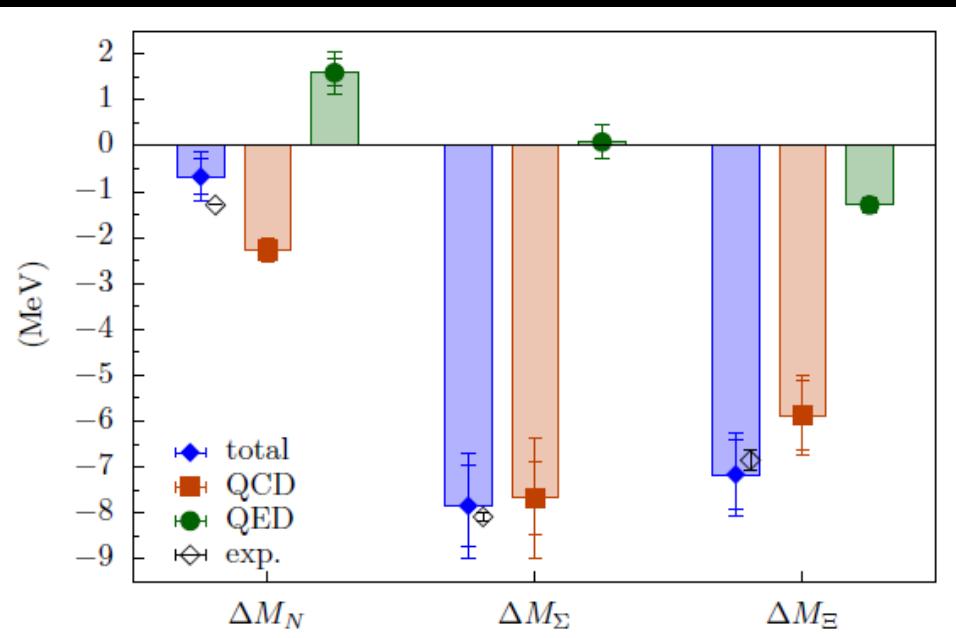
$m_{\pi} > 190$  MeV

# Hadron masses (2013)

$m_{\pi} = 135$  MeV with QED



BMW Collaboration, Science 322 (2008) 1224



BMW Coll.: Phys.Rev.Lett. 111 (2013) 252001

HPCI Program (FY2011-2015) Field 5: All Japan Computational Physics Collaboration

## The Origin of Matter and the Universe

(particle physics – nuclear physics – astrophysics, 11 institutions)



11 Pflops K Computer (RIKEN)



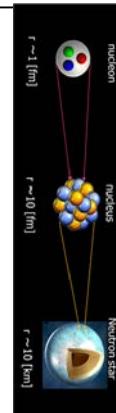
Lattice 2015 (July)  
Quark Matter 2015 (Oct.)

Project 1: Baryon-Baryon interaction from lattice QCD

Project 2: Nuclear quantum many-body calculation

Project 3: Supernova explosion and black-hole formation

Project 4: First stars and galaxies



Present un-physical point simulation  
for single and multi-baryons

On-going physical point simulation  
for single and multi-baryons in K

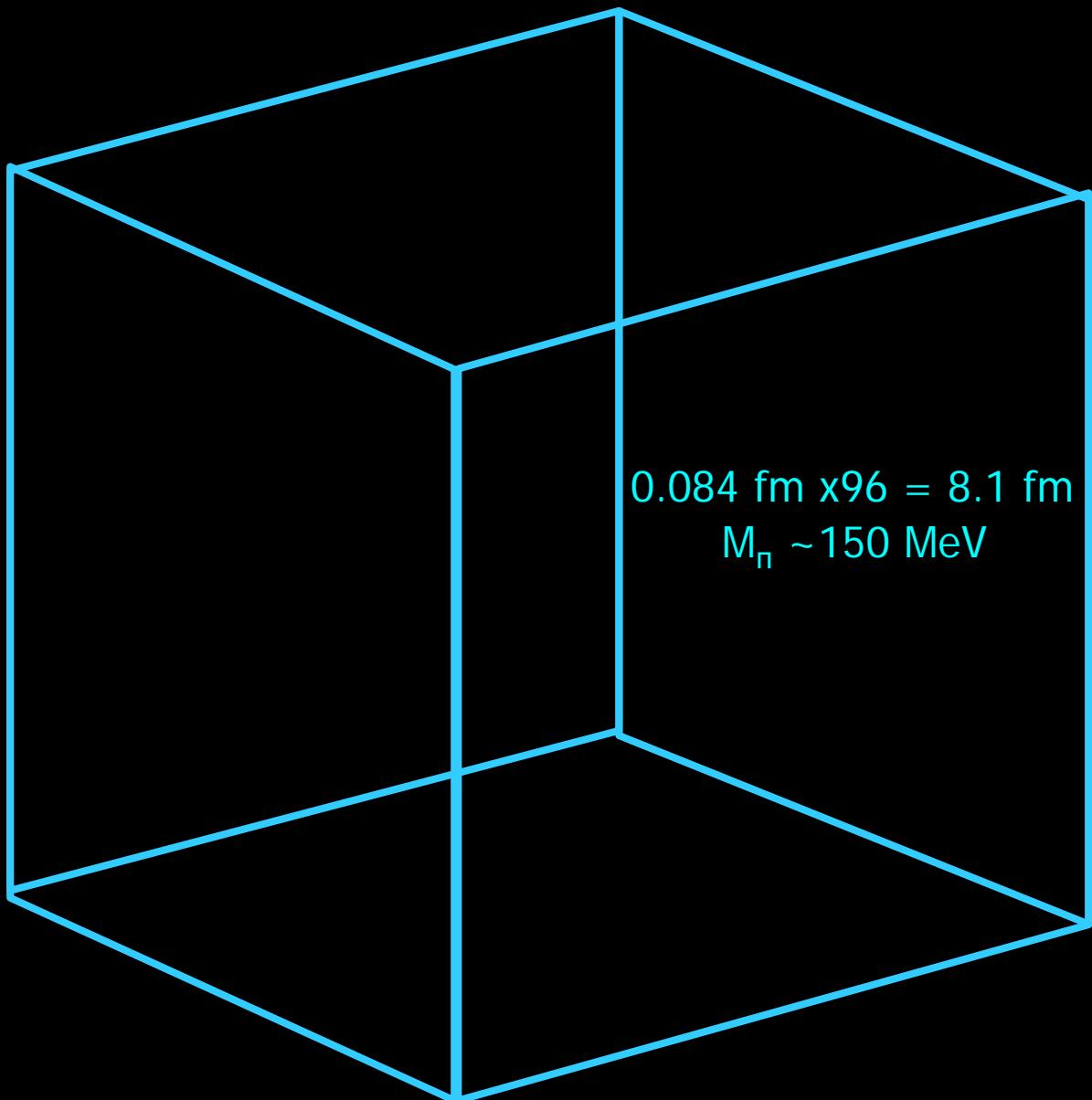
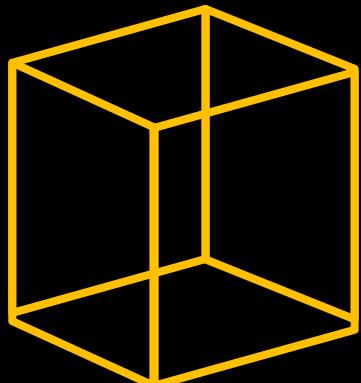
NEXT YEAR'S TALK  
Stay Tuned !!



TODAY'S TALK



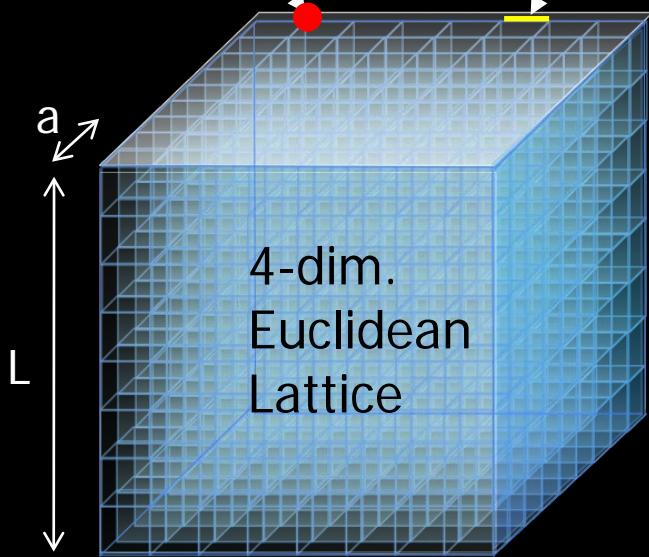
0.121 fm x32 = 3.9 fm  
 $m_\pi = 350-1200$  MeV



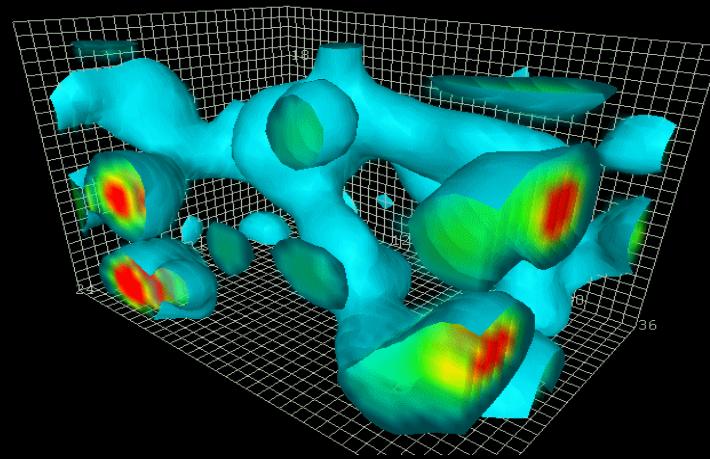
# 格子QCD



quarks  $q$   
on the sites      gluons  $U_\mu = \exp(iagA_\mu)$   
                        on the links



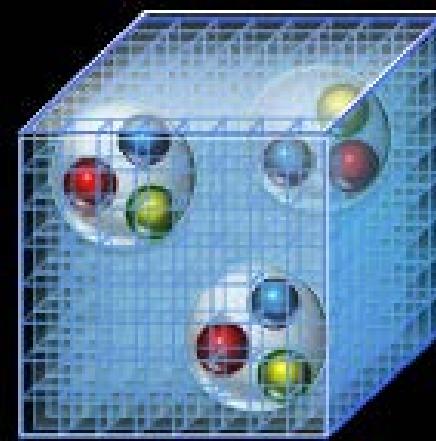
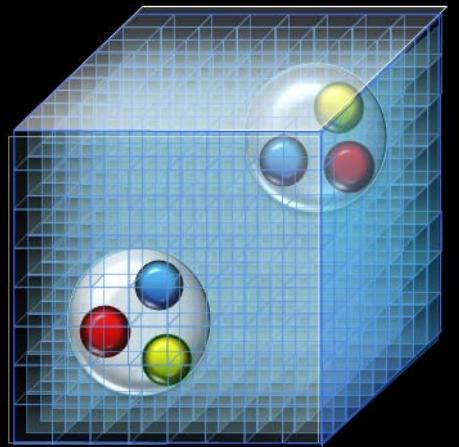
$$Z = \int [dU][dq d\bar{q}] \exp \left[ - \int d\tau d^3x \mathcal{L}_E \right]$$



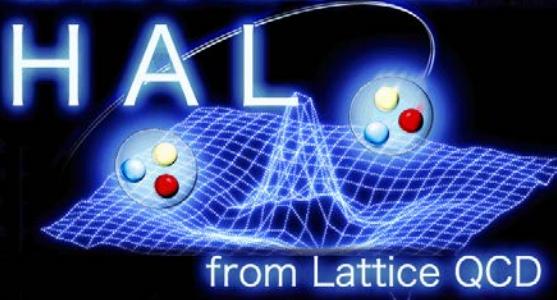
- well defined statistical system (finite  $a$  and  $L$ )
- gauge invariant
- fully non-perturbative

→ Numerical simulations

# Baryon Forces from LQCD



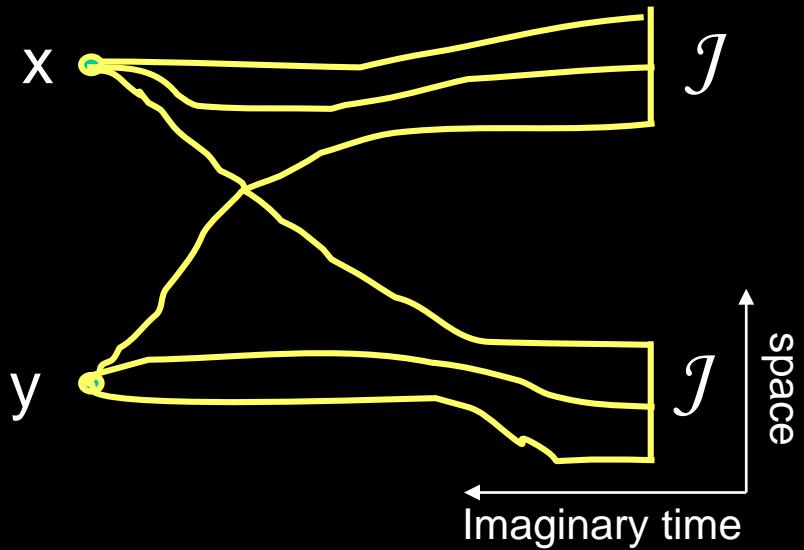
Hadrons to Atomic nuclei



|               |  |
|---------------|--|
| Univ. Tsukuba | T. Miyamoto, H. Nemura, K. Sasaki, M. Yamada |
| Univ. Tokyo   | B. Charron                                   |
| RIKEN         | T. Doi, T. Hatsuda, Y. Ikeda, V. Krejcirik   |
| Nihon Univ.   | T. Inoue                                     |
| YITP (Kyoto)  | S. Aoki, T. Iritani                          |
| RCNP (Osaka)  | N. Ishii, K. Murano                          |
| Birjand       | F. Etminan                                   |

Review: “Lattice QCD Approach to Nuclear Physics”  
HAL QCD Collaboration, Prog. Theor. Exp. Phys. 2012 (2012) 01A105

# Hadronic correlations in LQCD



$$\begin{aligned} \langle N_1(\mathbf{x}, t) N_2(\mathbf{y}, t) \mathcal{J}_1^\dagger(0) \mathcal{J}_2^\dagger(0) \rangle \\ = \sum_n \langle 0 | N_1(\mathbf{x}) N_2(\mathbf{y}) | n \rangle a_n e^{-E_n t} \\ \xrightarrow{t > t^*} \phi(\mathbf{r}, t) = \sum_{n < n^*} b_n \phi_n(\mathbf{r}) e^{-E_n t} \end{aligned}$$

Finite Volume Method :

$E_0(L) \rightarrow$  phase shift

Luescher, Nucl. Phys. B354 (1991) 531

HAL QCD Method

$\phi(r,t) \rightarrow$  kernel  $\rightarrow$  phase shift

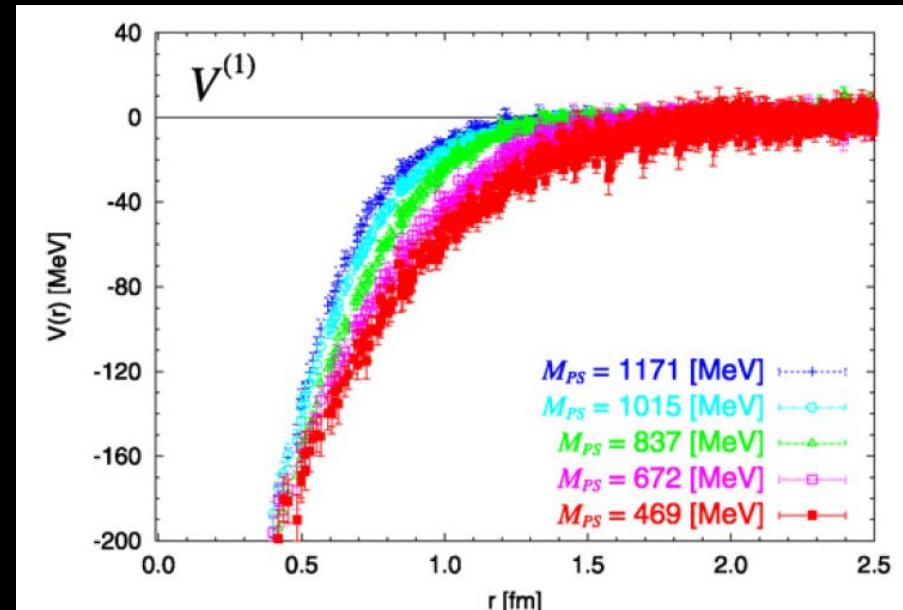
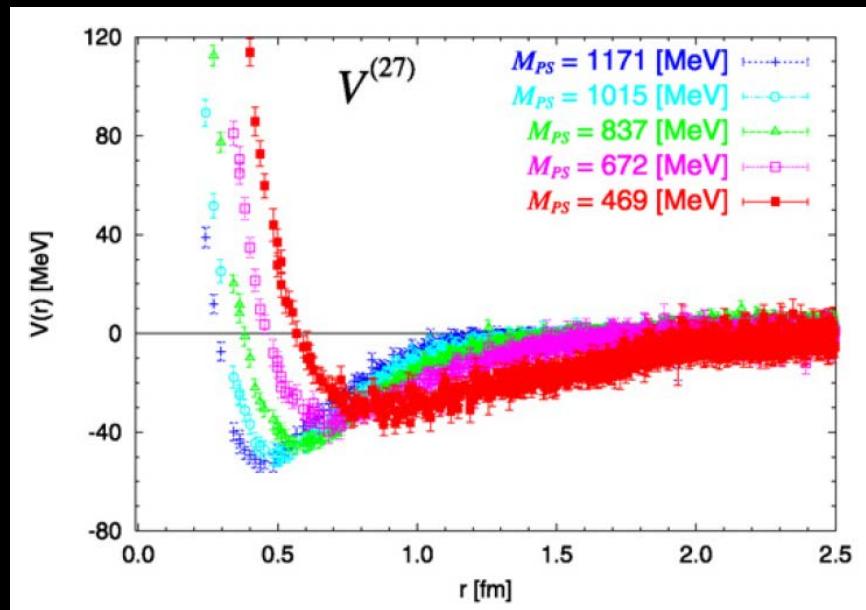
Ishii, Aoki & Hatsuda, PRL 99 (2007) 022001  
 Comp. Sci. Dis. 1 (2009) 015009  
 HAL QCD Coll., PLB 712 (2012) 437

# BB Forces in 3-flavor QCD

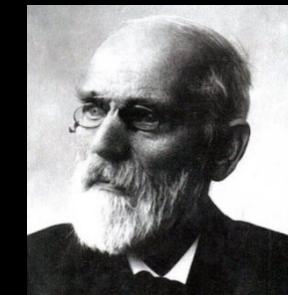
HAL QCD Coll.  
Phys. Rev. Lett. 106 (2011) 162002  
Nucl. Phys. A881 (2012) 28

PP (uud-uud) channel  
(partial) Pauli blocking

H (uds-uds) channel  
No Pauli blocking



Pauli and van der Waarls  
at work !



# BB Forces in 3-flavor QCD

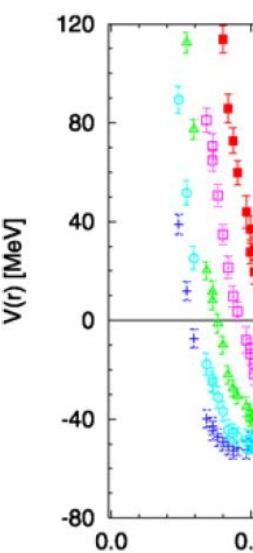
HAL QCD Coll.

Phys. Rev. Lett. 106 (2011) 162002

Nucl. Phys. A881 (2012) 28

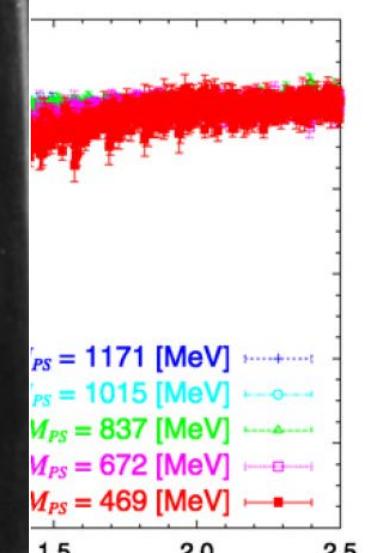
PP ( $\psi\bar{\psi}$ - $\psi\bar{\psi}$ ) channel

(p)



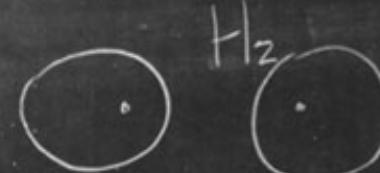
UU ( $u\bar{u}d\bar{d}$ - $u\bar{u}d\bar{d}$ ) channel

cking

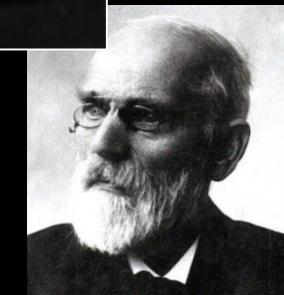


$$V = \sum P^a V_P$$

$$V = V_0 - \sum V_{r_A} \left\{ 1 - \frac{1}{(r_A)^2} \right\}$$



Pauli and van der Waals  
at work !

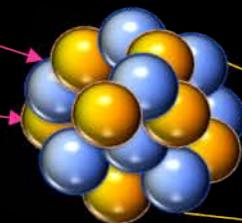


# Nuclear Matter, Neutron Matter & Finite Nuclei from LQCD + BHF

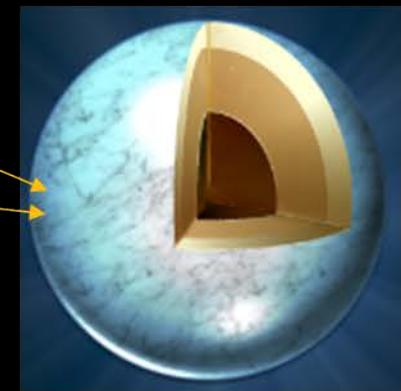
nucleon  $\sim 1$  [fm]



nucleus  $\sim 10$  [fm]



Neutron star  $\sim 10$  [km]

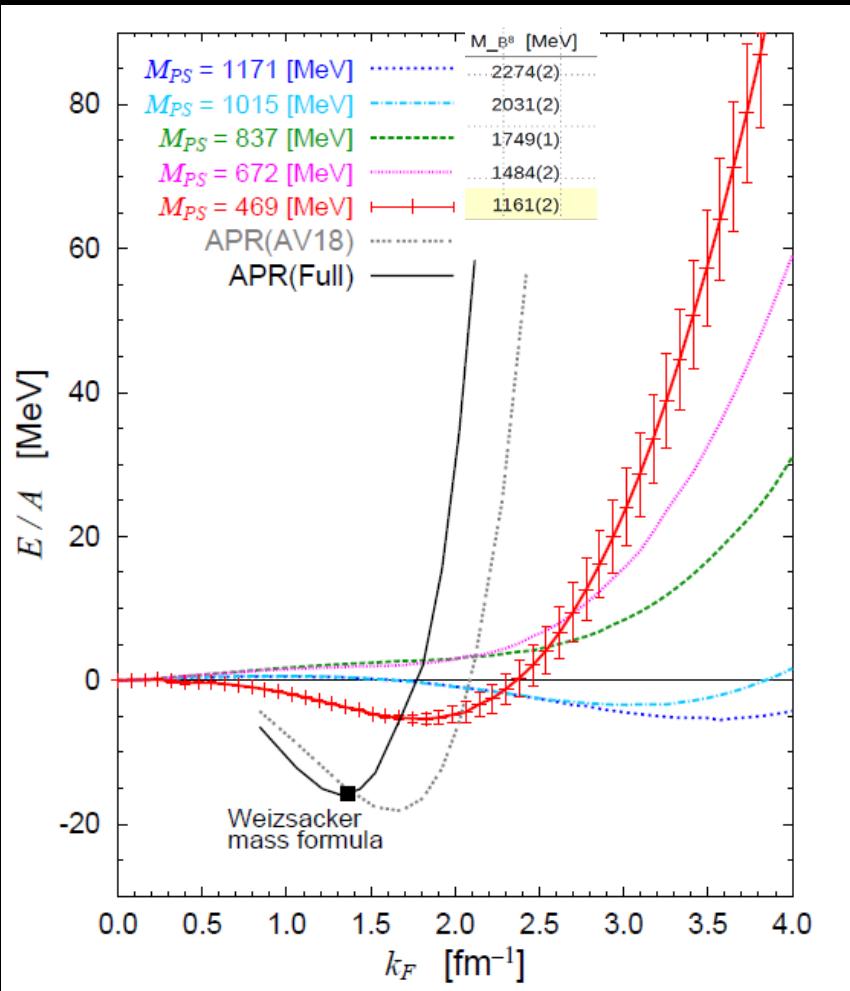


# Nuclear EOS from Lattice NN force + BHF calculation

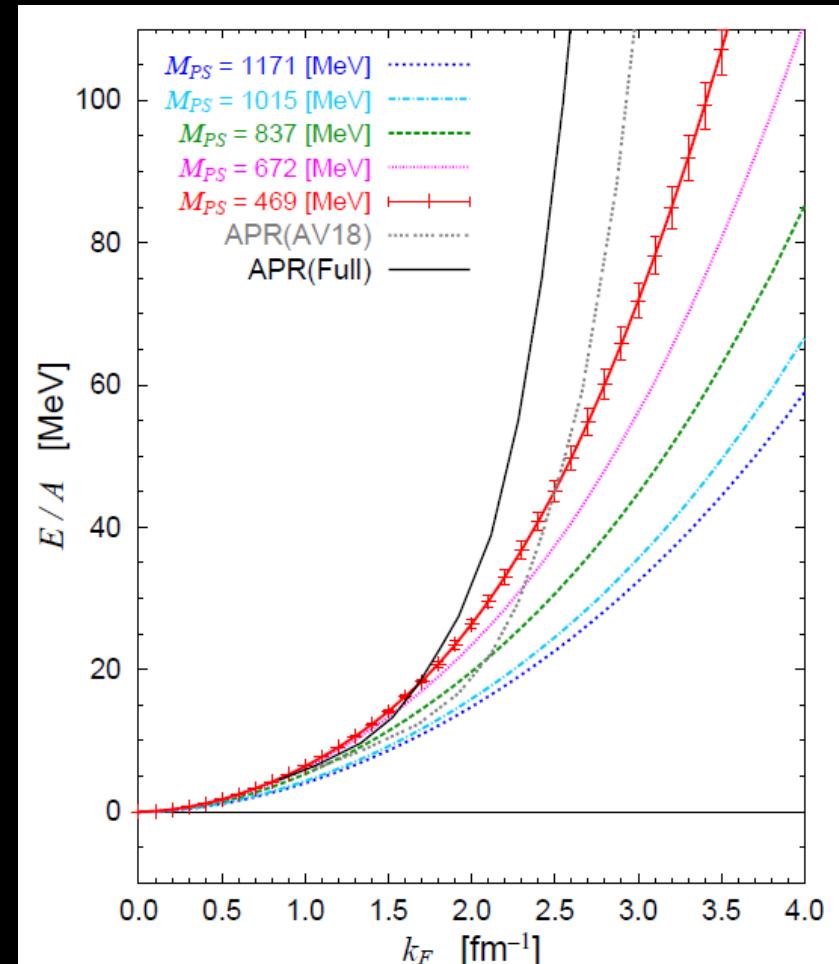
(NN force:  $^1S_0$ ,  $^3S_1$ ,  $^3D_1$  channels only)

HAL QCD Coll., Phys. Rev. Lett. 111 (2013) 112503

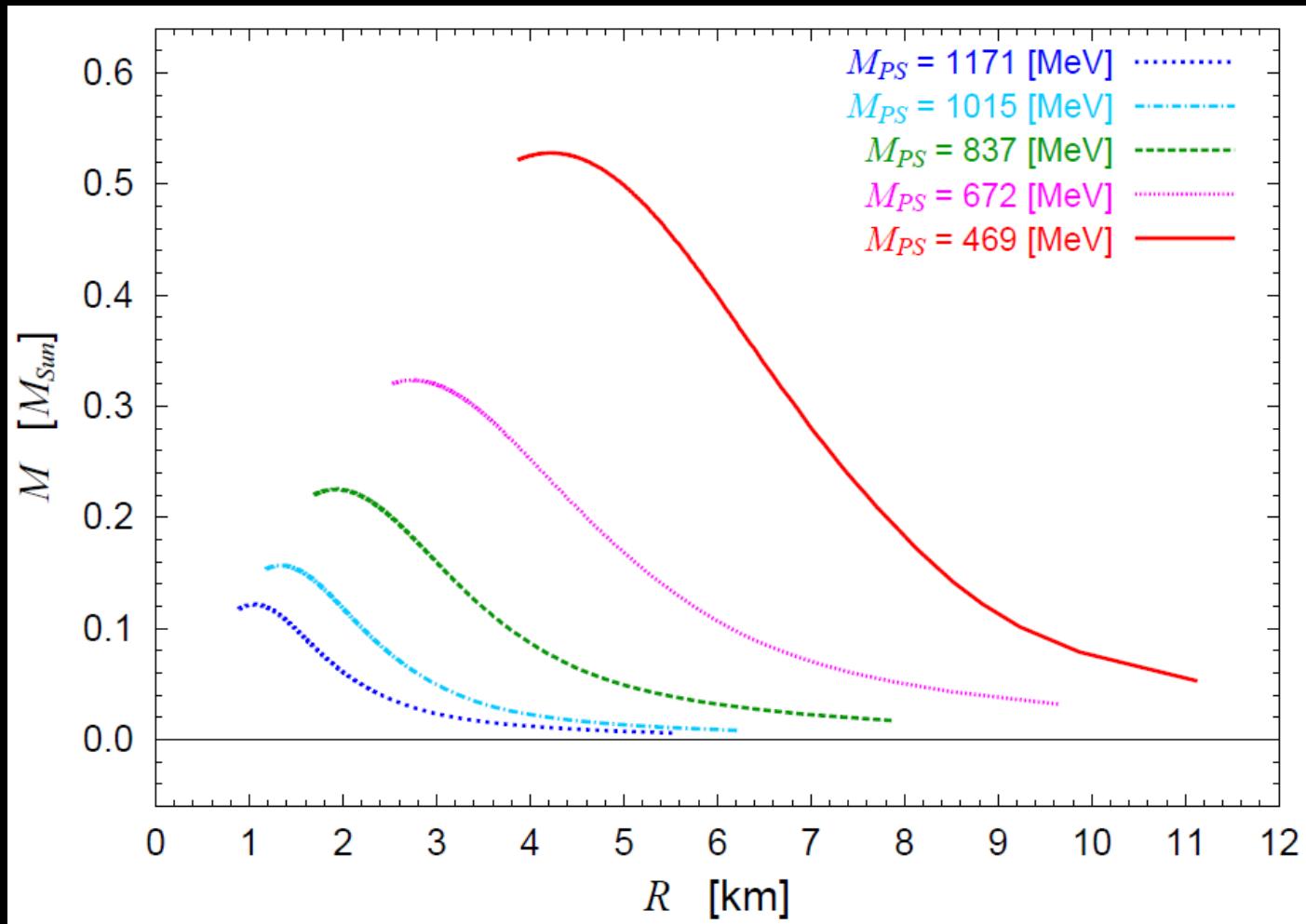
## Nuclear Matter



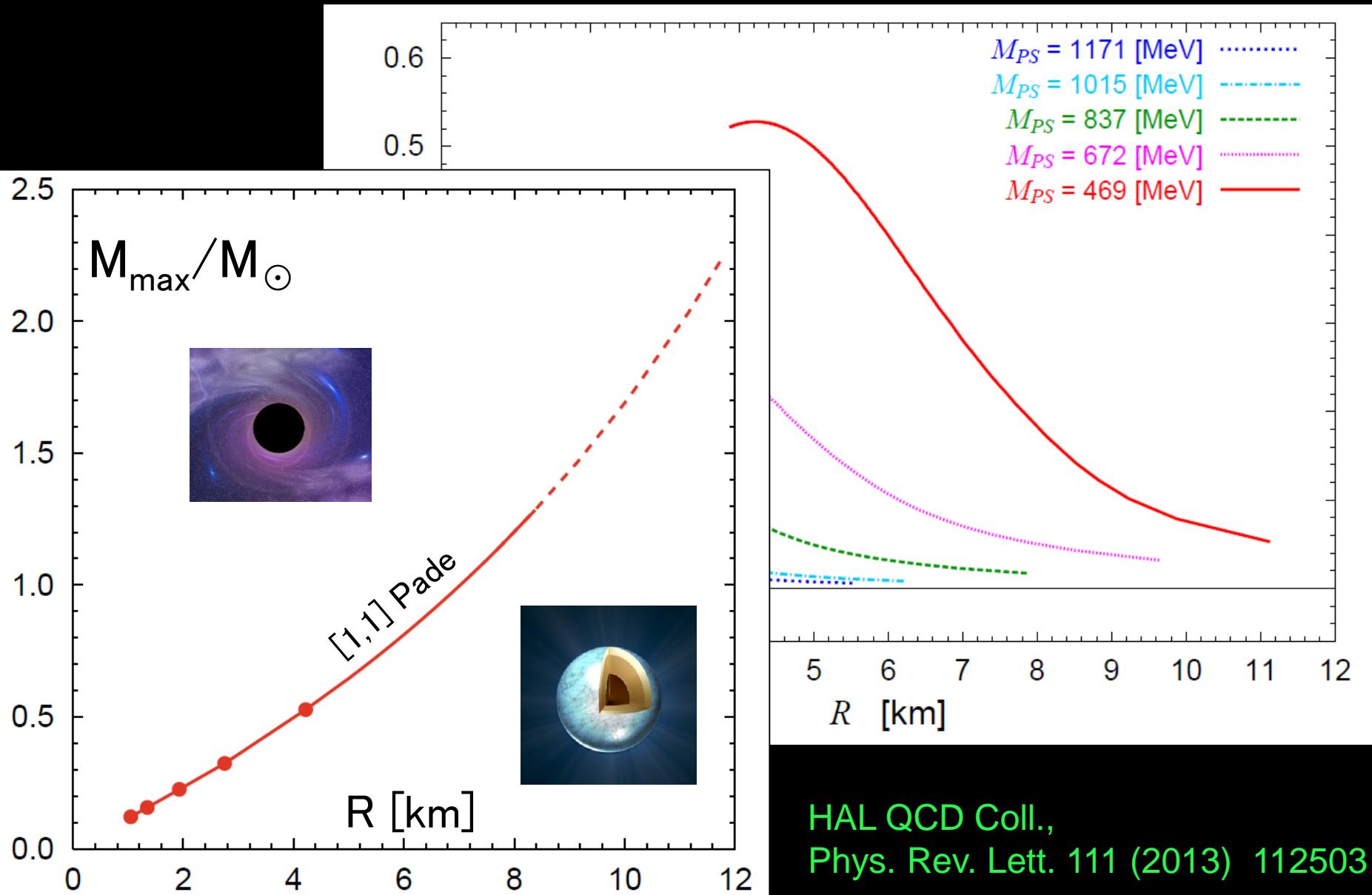
## Neutron Matter



# Neutron Star from “Lattice EOS”



# Neutron Star from “Lattice EOS”

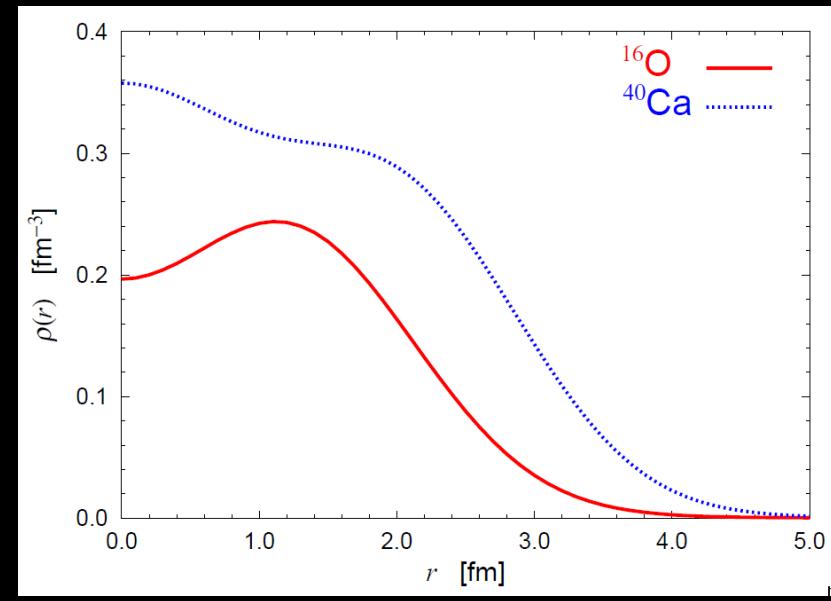
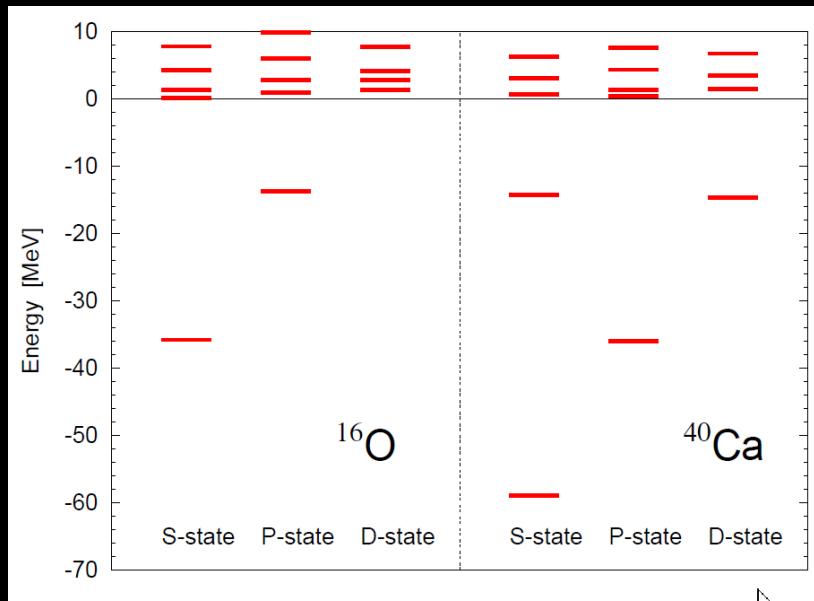


# Finite Nuclei from Lattice NN force + BHF calculation

(NN force:  $^1S_0$ ,  $^3S_1$ ,  $^3D_1$  channels only)

Inoue et al. [HAL QCD Coll.], arXive 1408.4892

Bound nuclei start to appear from  $m_\pi = 470$  MeV

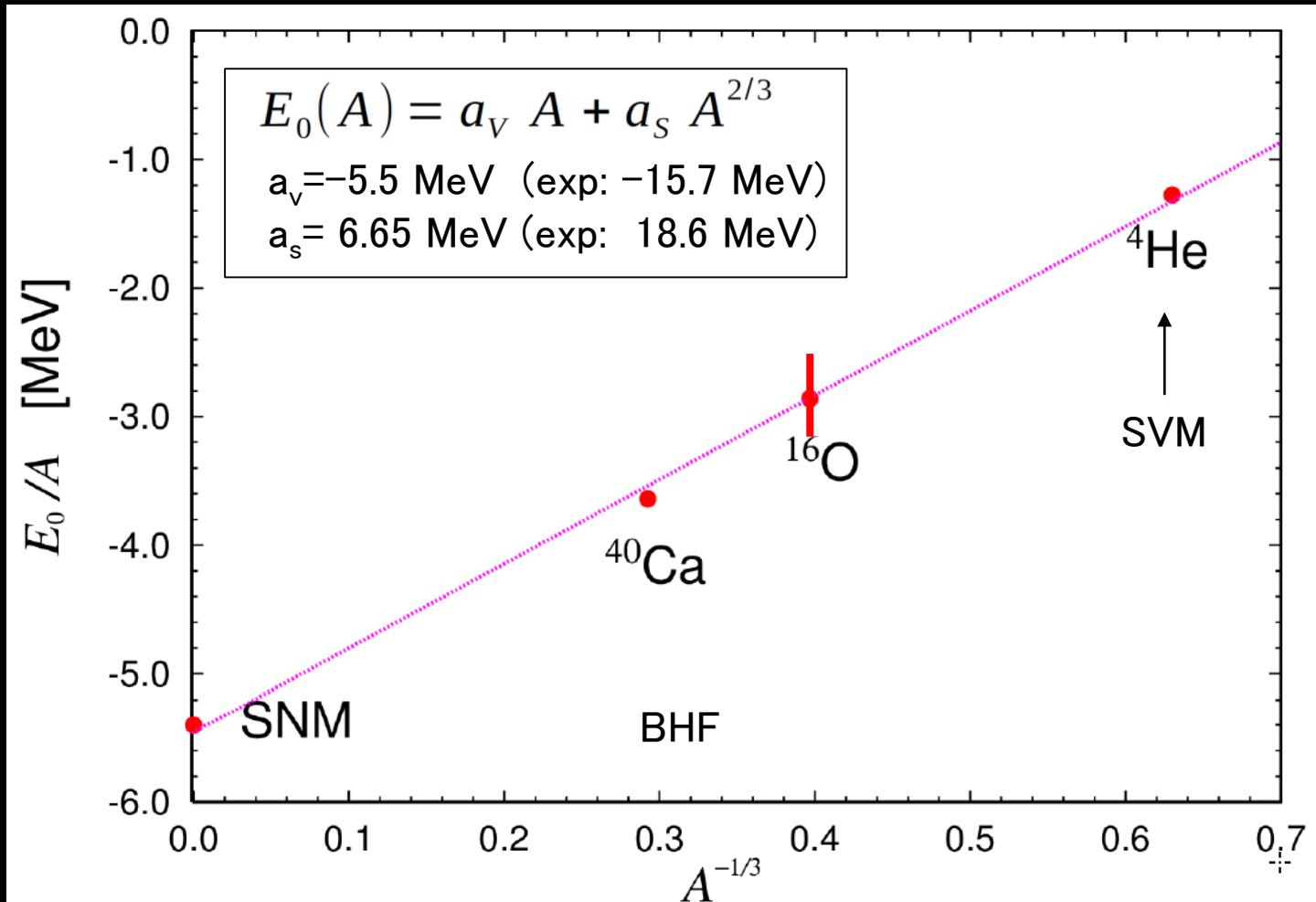


|                  | Single particle level |       |       |       | Total energy |         | Radius                       |
|------------------|-----------------------|-------|-------|-------|--------------|---------|------------------------------|
|                  | $1S$                  | $1P$  | $2S$  | $1D$  | $E_0$        | $E_0/A$ | $\sqrt{\langle r^2 \rangle}$ |
| $^{16}\text{O}$  | -35.8                 | -13.8 |       |       | -34.7        | -2.17   | 2.35                         |
| $^{40}\text{Ca}$ | -59.0                 | -36.0 | -14.7 | -14.3 | -112.7       | -2.82   | 2.78                         |

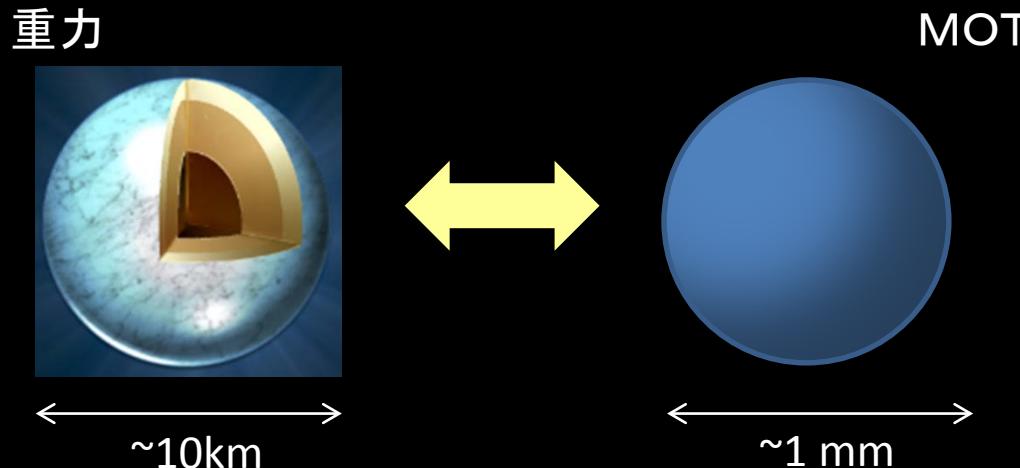
# Nuclear Binding Energy from Lattice NN Force

Inoue et al. [HAL QCD Coll.], arXive 1408.4892

Bethe–Weizacker behavior at  $m_\pi = 470$  MeV



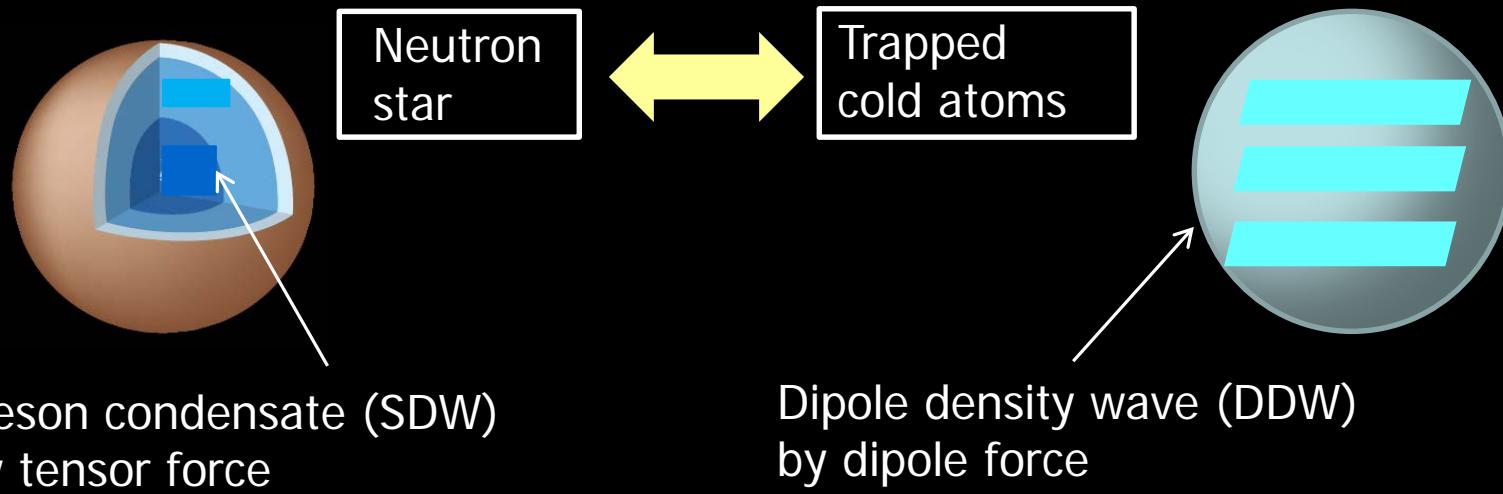
# 冷却原子を用いた 中性子星の実験室シミュレーション？



- ハドロン相とクオーク相のクロスオーバー  
 $\Leftrightarrow$  冷却 ボース原子-冷却フェルミ原子混合気体  
Maeda, Baym & Hatsuda, Phys. Rev. Lett. 103 (2009) 085301
- $\pi$ 中間子凝縮  $\Leftrightarrow$  冷却双極子原子(分子)気体  
Meada, Baym & Hatsuda, Phys. Rev. A 87 (2013) 021604(R)

# meson-condensation in ultracold dipolar atoms

Maeda, Baym, Hatsuda  
(2012)

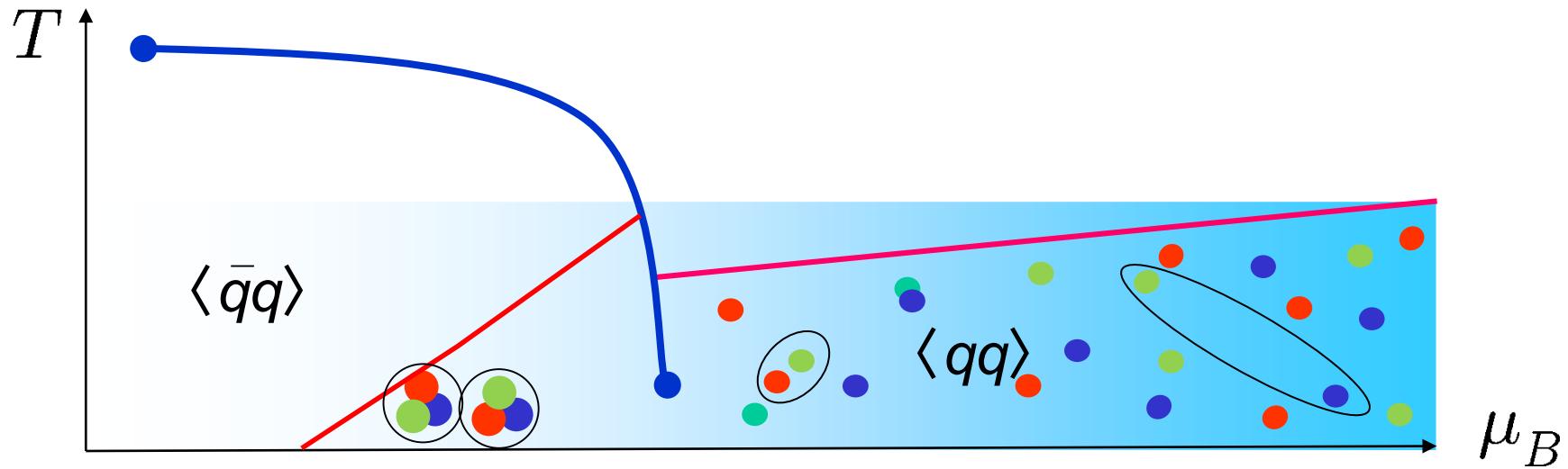


The equation shows the dipole-dipole interaction potential  $U_{dd}$  between two dipoles:

$$U_{dd} = -\frac{3(\vec{\mu}_1 \cdot \hat{r})(\vec{\mu}_2 \cdot \hat{r}) - \vec{\mu}_1 \cdot \vec{\mu}_2}{r^3} - \frac{8\pi}{3}\vec{\mu}_1 \cdot \vec{\mu}_2 \delta(\vec{r})$$

| neutron matter                    | atomic dipoles  |
|-----------------------------------|---|
| neutron stars                     | atomic gases(e.g. $^{163}\text{Dy}$ , $^{161}\text{Dy}$ ) |
| neutrons                          | fermionic dipoles   |
| spin                              | pseudo-spin (hyperfine states)                            |
| neutral vector meson ( $\rho^0$ ) | photon  |
| tensor-force potential            | dipole-dipole interaction potential                       |
| meson condensation                | spin/density wave   |

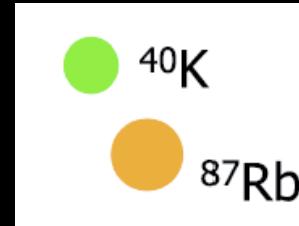
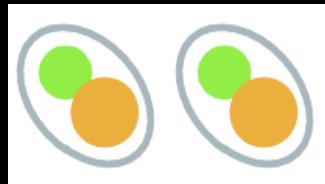
# Hadron-Quark Phase Transition and the role of “Diquarks”



Nuclear superfluid  $\leftrightarrow$  Fermion+Diquark  $\leftrightarrow$  Diquark superfluid



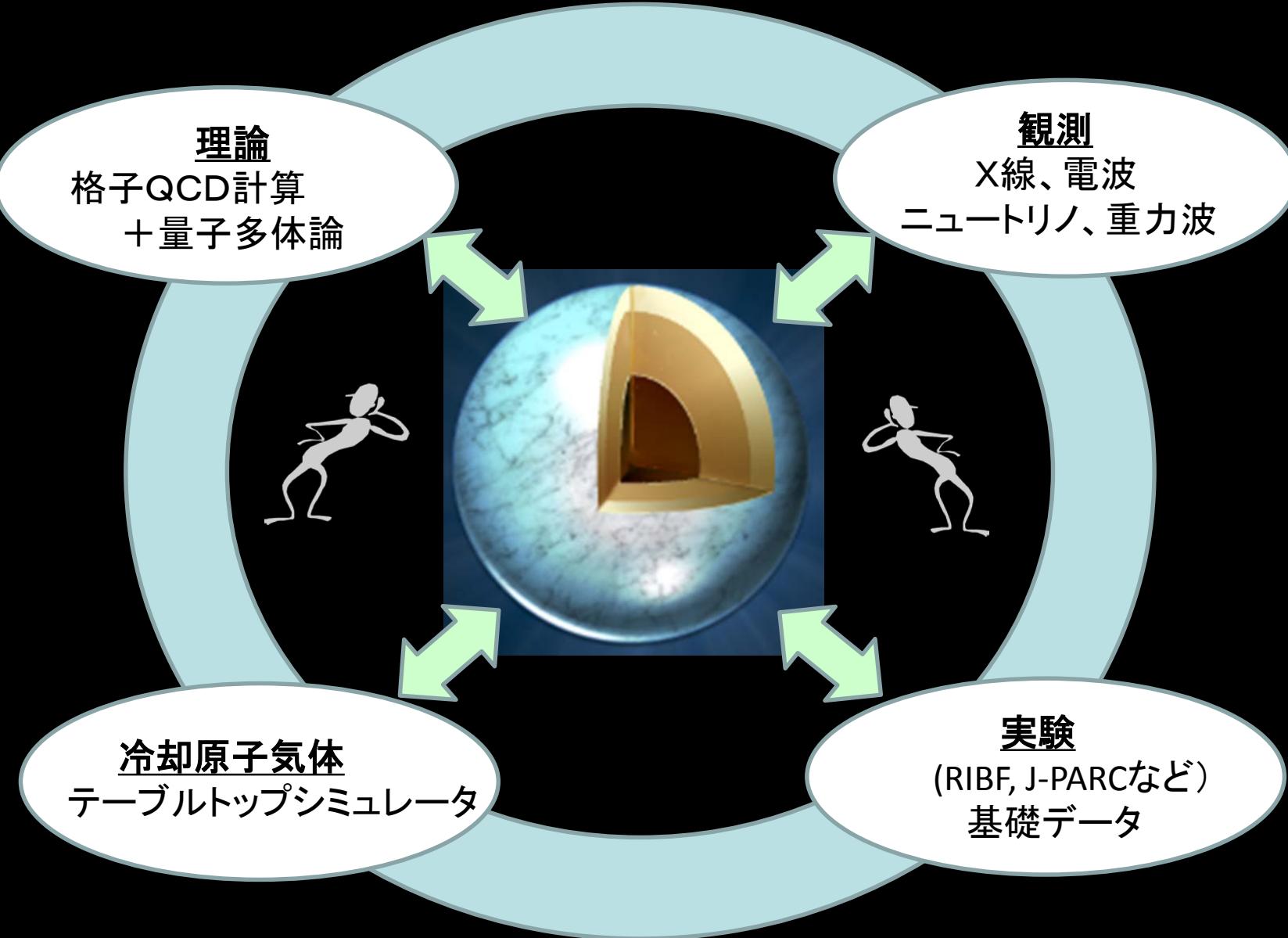
Induced superfluid  $\leftrightarrow$  Fermi-Bose mixture



$$a_{NN}^{\text{Born}} = -\frac{m_N}{2m_R} a_{bf}$$

Maeda, Baym & Hatsuda,  
Phys. Rev. Lett. 103 ('09)

# まとめ



Present un-physical point simulation  
for single and multi-baryons

On-going physical point simulation  
for single and multi-baryons in K

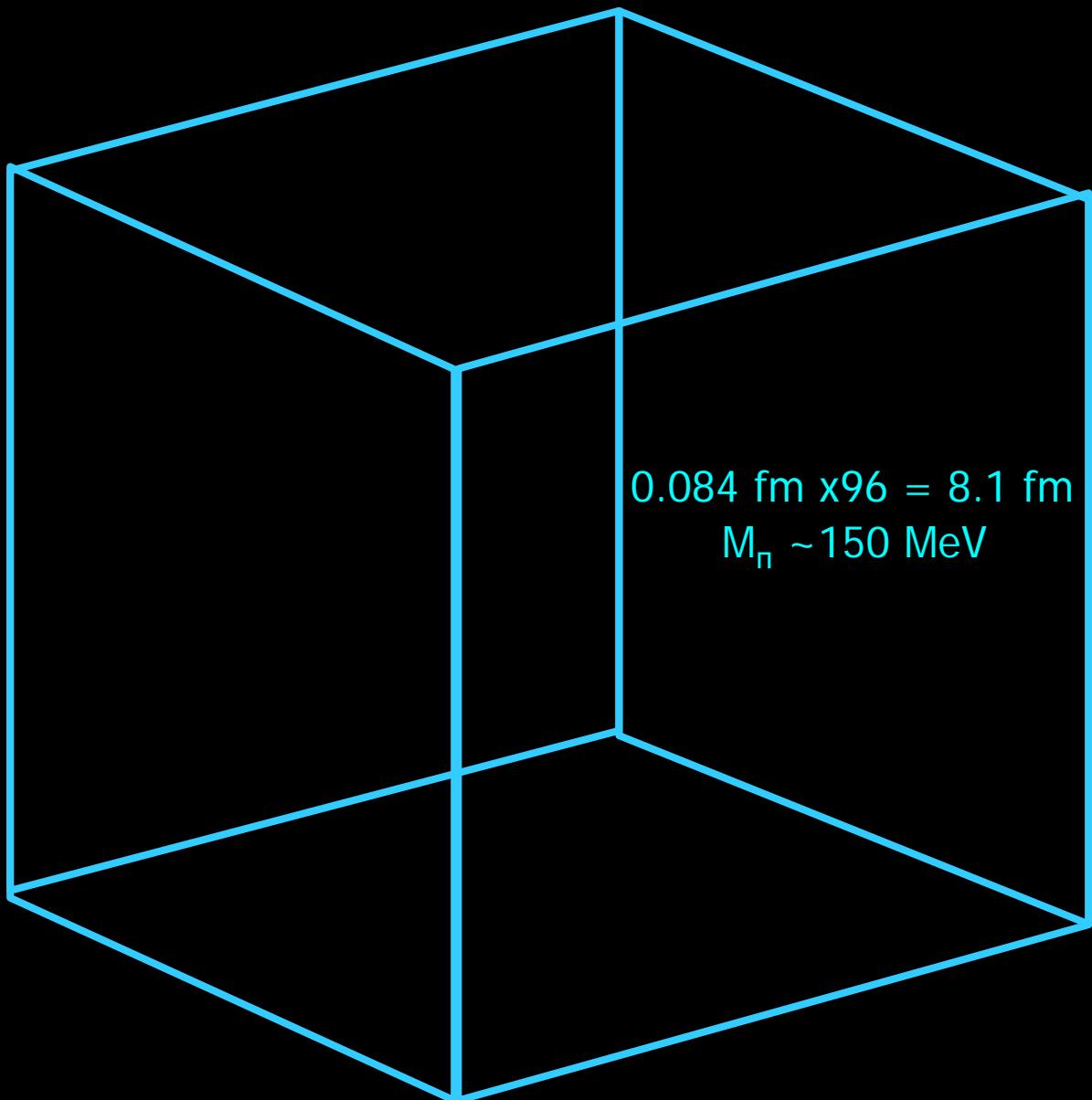
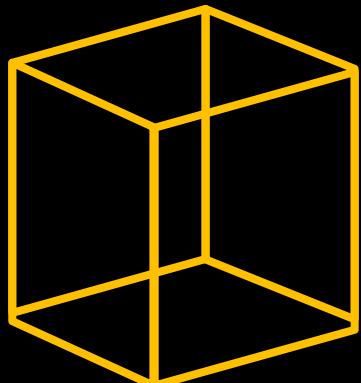
NEXT YEAR'S TALK  
Stay Tuned !!



TODAY'S TALK



0.121 fm x32 = 3.9 fm  
 $m_\pi = 350-1200$  MeV



END